Poster

ANSWERS TO SOME IMPORTANT QUESTIONS ON THE USE OF GR MODEL

E. Lyratzi^{1,2}, D. Stathopoulos^{1,2}, E. Danezis¹, A. Antoniou¹ and D. Tzimeas¹

 ¹ University of Athens, Faculty of Physics, Department of Astronomy, Astrophysics and Mechanics, Panepistimioupoli Zographou, GR 15784 Athens, Greece
² Eugenides Foundation, Syngrou 387, 175 64 P. Faliro, Greece

E-mail: elyratzi@phys.uoa.gr, dstatho@phys.uoa.gr, edanezis@phys.uoa.gr, ananton@phys.uoa.gr, dtzimeas@phys.uoa.gr

After a decade of using the GR model (Gauss-Rotation model) many interesting questions have been arisen concerning the physics and the mathematical expression of the model, as well as the exact number of components that are required to reproduce the complex absorption and emission profiles and the uniqueness of the calculated values of the parameters (kinematical parameters, the optical depth the Full Width at Half Maximum (FWHM), the column density and the absorbed or emitted energy), which are calculated with this method. In this paper we try to answer some of these basic questions.

Poster

ON THE STARK BROADENING OF Lu III SPECTRAL LINES

Z. Majlinger¹, Z. Simić² and M. S. Dimitrijević²

 ¹Faculty of Humanities and Social Sciences, University of Rijeka, 51000 Rijeka, Croatia
²Astronomical Observatory, Volgina 7, 11060 Belgrade, Serbia

E-mail: zlatko.majlinger@gmail.com, zsimic@aob.bg.ac.rs, mdimitrijevic@aob.bg.ac.rs

The electron-impact widths for six Lu III spectral lines have been calculated by using the modified semiempirical method. With obtained results, the influence of Stark broadening on Lu III lines has been investigated in the spectra of A-type stars.