

MOL-D

A COLLISIONAL DATABASE AND WEB SERVICE WITHIN THE VIRTUAL ATOMIC AND MOLECULAR DATA CENTER

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The data

- Photo-dissociation cross-sections for the individual ro-vibrational states (a pair of quantum numbers J and v) of diatomic molecular ions
- Ro-vibrational energy for the corresponding state
- \bullet 423 states for $H_2^{\ +}$, 833 for $He_2^{\ +}$
- Data for 25 discrete wavelengths (A):

50 60 70 80 90 100 200 300 400 500 600 700 800 900 1000 1100 1200 1300 1400 1500 1600 1700 1800 1900 2000



The physics

Theoretical calculations
 (photon + molecular ion =
 atom + ion)

$$h\nu + A_2^+ \iff A + A^+$$

- The cross-sections for the photodissociation of individual ro-vibration state of the considered molecular ions are determined in the dipole approximation
 - individual ro-vibration state of the considered molecular $\sigma_{J,v}(\lambda) = \frac{8\pi^3}{3\lambda} \left[\frac{(J+1)|D_{E,J+1;v,J}|^2 + J|D_{E,J-1;v,J}|^2}{2J+1} \right]$
- Mihajlov and coworkers
 (Mihajlov & Dimitrijevic
 1986; Mihajlov et al. 1993,
 2007; Ignjatovic et al. 2009,
 2014b; Sreckovic et al. 2014)



Physics (2)

 Corresponding average thermal crosssections are given by:

$$\sigma_{\rm ph}(\lambda, T) = \frac{1}{Z} \sum_{J} \sum_{v} g_{J;v} (2J+1) e^{-\frac{E_{Jv} - E_{00}}{k_{\rm B}T}} \sigma_{J,v}(\lambda)$$



Importance

 Hydrogen and helium molecular ion data are important for calculation of solar and stellar atmosphere models and for radiative transport, as well as for kinetics of other astrophysical and laboratory plasma (i.e. early Universe).



How to get from this...

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						0.285062E-17 0.242410E-17								
						0.996591E-18								
						0.799153E-19								
						0.129540E-18								
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						0.717645E-18 0.381651E-18								
						0.132813E-18								
						0.158131E-19								
						0.568653E-20								
0	168834E-21	0.247946E-1	8 0.456043E-18	3 0.471815E-1	9 0.414585E-18	0.524763E-19	0.544617E-18	0.356561E-18	0.895826E-18	0.140265E-17	0.165285E-18	0.215277E-17	0.182719E-17	0.357944E-17
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						0.672568E-19								
						0.259097E-19								
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...to this...

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...and why?

- Fully annotated, interconnected data
- Standardized representations of species, states and their properties
- Metadata on scientific sources
- Machine readable format (human too if needed)
- Apply XSL transformations to any format (HTML, SME...)
- Convert to objects
- Interoperability

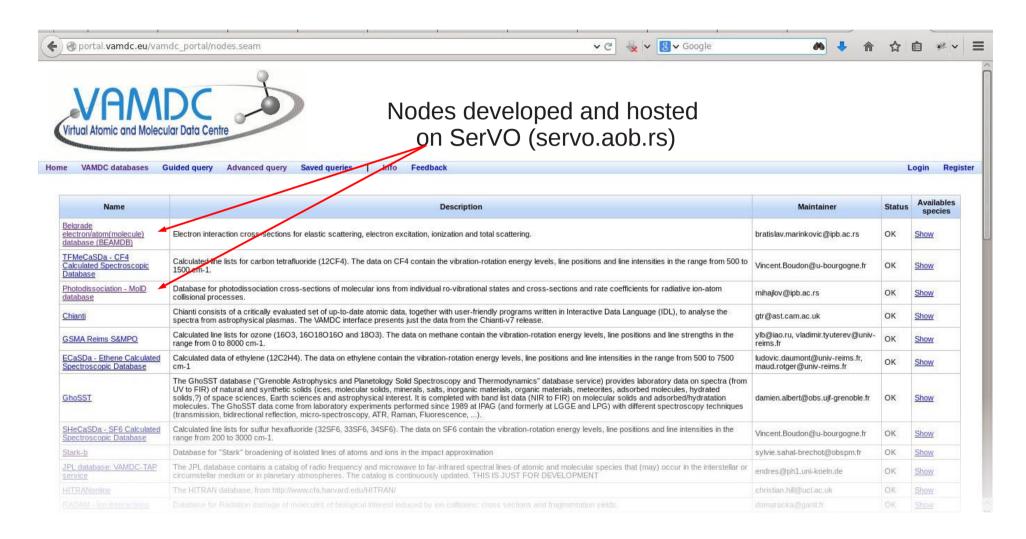


VAMDC philosophy

- Local providers expose data through standardized web services ('nodes')
- Central repository holds list of nodes and executes distributed, aggregate queries.
- Available as a web portal / SOAP web service / Astrogrid VODesktop / Java library
- Extensions of IVOA protocols (VAMDC-TAP), XML standard for data representation (XSAMS)
- Transform data for further analysis ("processors")

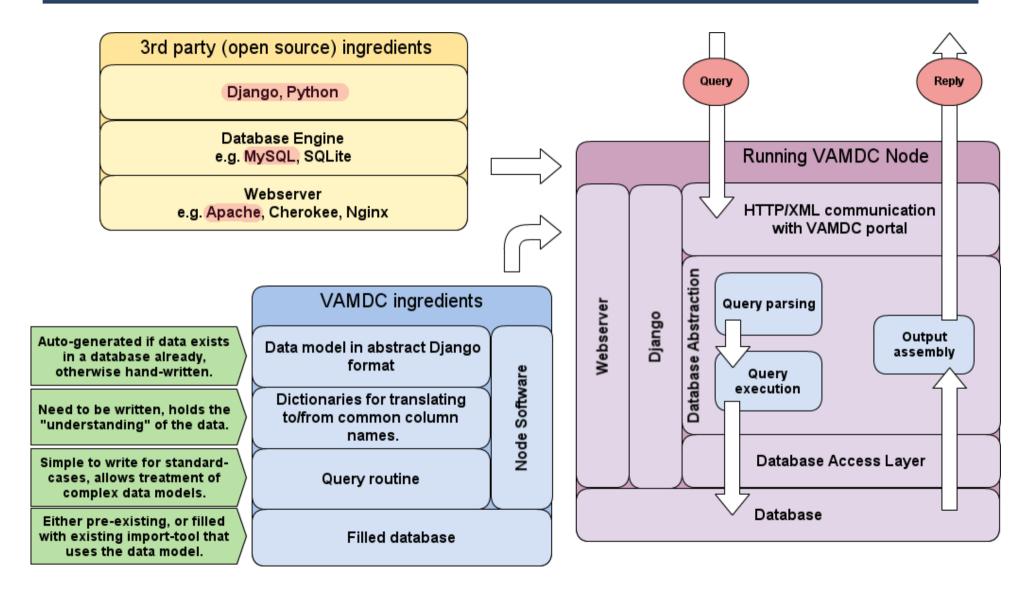


VAMDC Nodes





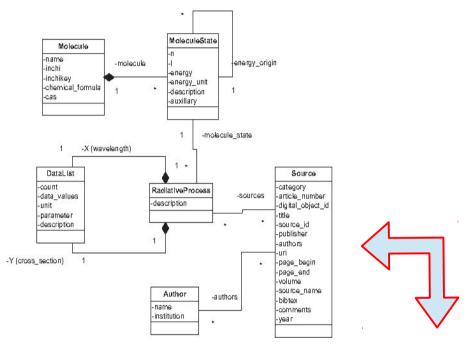
VAMDC NodeSoftware





Django is convenient

- Write static structure (models.py), django generates db for you
- OR
- Generate models from existing db
- Decide how to describe your data in XSAMS schema (dictionaries.py)
- Map models to XSAMS via queryfunc.py (it's ok for models to be incompatible with XSAMS schema)

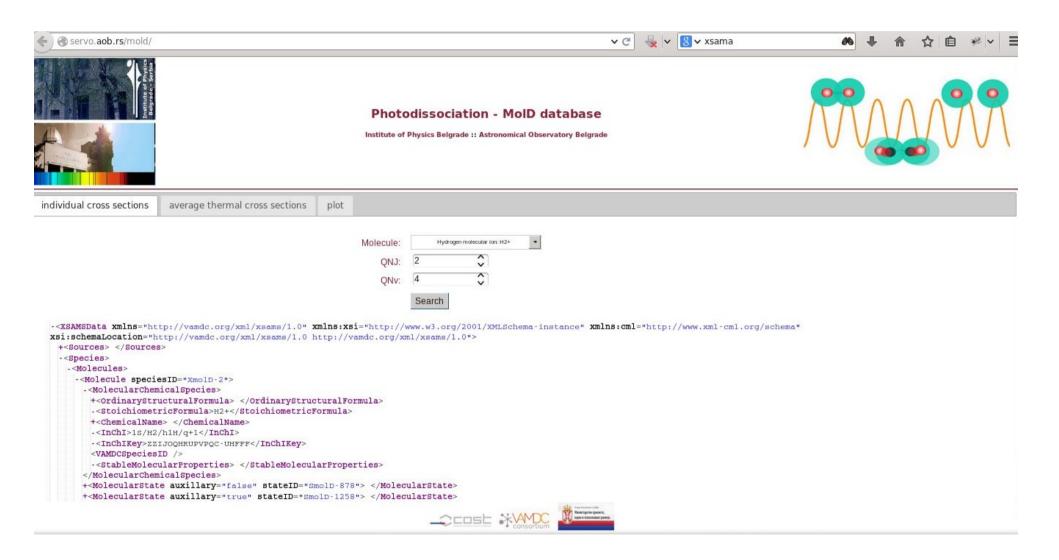


web service & web interface

- http://servo.aob.rs/mold
- VAMDC asks only for a standardcompliant and validated web service
- Build more stuff on top of Django/Python
- Easy template / url handling
- AJAX enabled interface

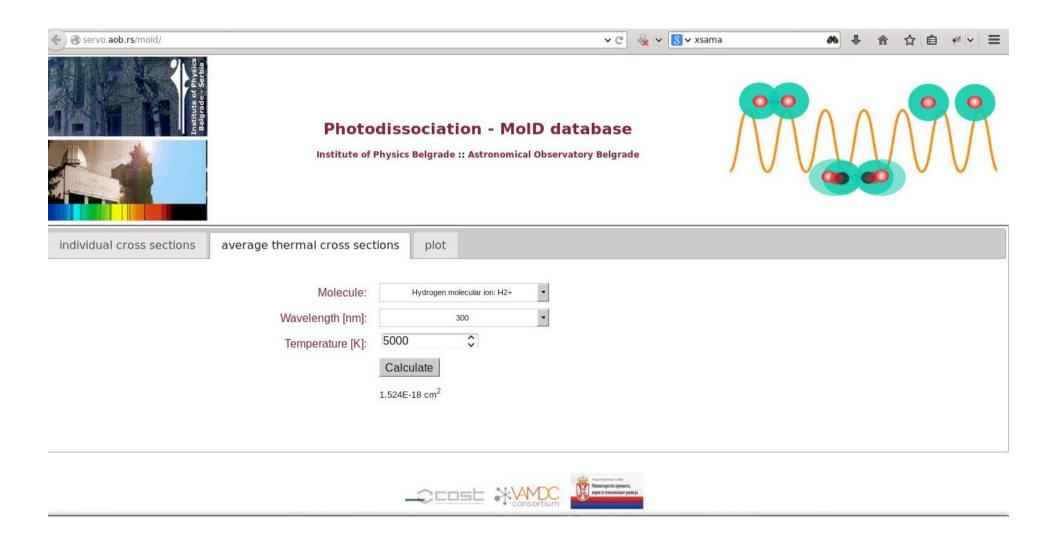


xsams



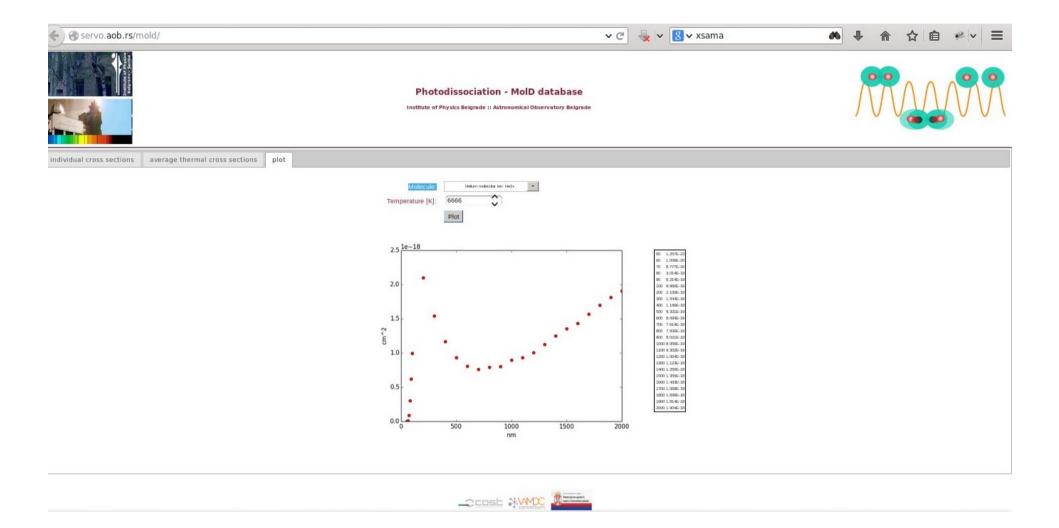


avg thermal cs's





plot





Further development

- Include new data about processes which involve species such as HeH⁺, LiH⁺, NaH⁺, SiH⁺ which are important for the early universe chemistry, the stellar and solar atmosphere modeling.
- Results of the rate coefficients for the ion-atom absorption processes and inverse emission processes
- Inclusion of the rate coefficients for the chemi-ionization in atom-Rydberg atom collisions (including the processes of the associative and Penning type ionization) and corresponding inverse chemi-recombination processes in electron-ion-atom collisions.