

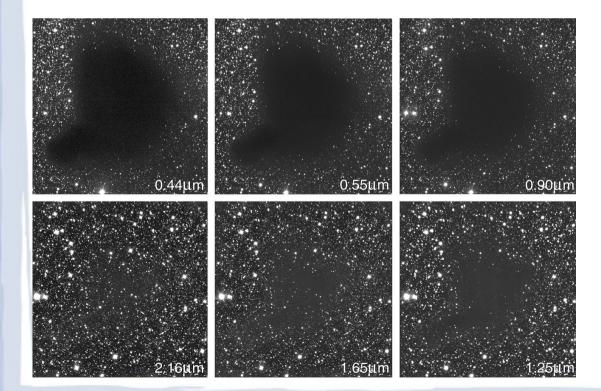
Faculty of Mathematics University of Belgrade

Atomic and Molecular Data - Application on Formation of Molecules in Dark Clouds

Dobardžić. A. & Kovačević, A.

Astrochemistry

 The study of the synthesis of molecules in space and their role in determining properties of Interstellar Medium.



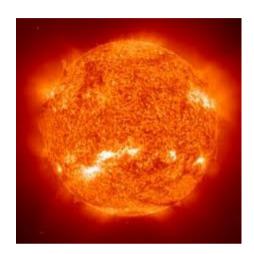
Top row – optical images of B68

Bottom row – IR images of B68 (Dust extinction is less

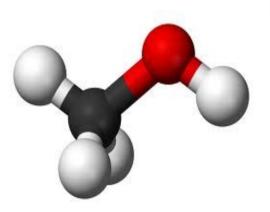
efficient at longer wavelengths)

Interstellar medium (ISM)

- 20 30 % of our Galaxy
- T = 10 10⁴ K (Sun temperature 5778 K)
- ρ = 100 10⁸ H atoms cm⁻³



- H is most abundant; He ~ 10%; C,N,O ~ 0.1%
- 1% of ISM mass is in grains
- Organic and inorganic molecules
- Complex chemistry



Dark clouds

- T ~ 10 K
- $\rho = 10^{10-12} \text{ H atoms m}^{-3}$

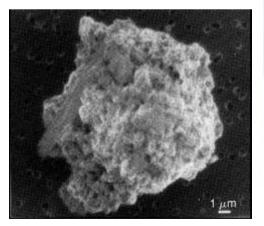


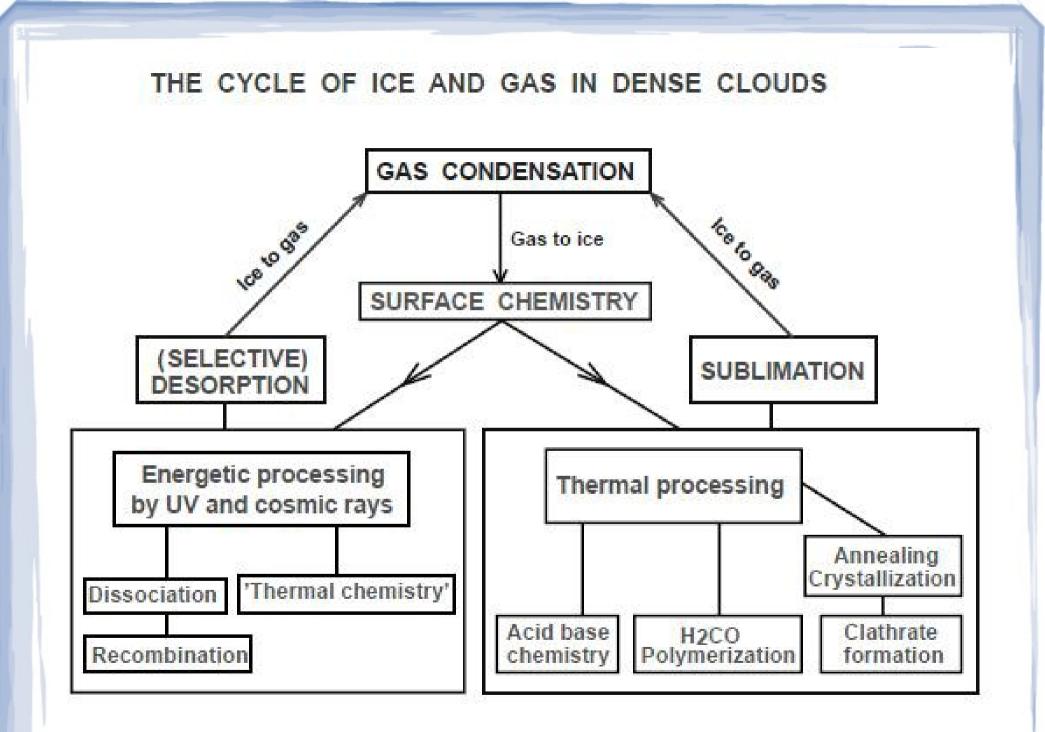
- Cold gas chemistry+ grain chemistry
- Not penetrated by optical and UV photons
- Little ionisation
- Interesting because material is mostly molecular, dominant species is H₂.



Dark clouds

- Cold gas and grain chemistry
- CO, N₂, O₂, C₂H₂, HCN, simple carbon chains
- Grain chemistry H_2 , CO_2 , CH_3OH
- Largest detected molecule HC₁₁N
- 1-500 solar masses
- 1-5 pc (distance to Proxima Centauri 1.3 pc)
- Can form one or a couple of low-mass stars





and the second se

Chemical Kinetics

(**first order** - rate of formation and loss proportional to the concentration of one reactant)

 $\begin{array}{ll} A+hv \rightarrow C+D & \beta \ (units \ s^{-1}) \end{array}$ Loss of A per unit volume per unit second is: dn(A)/dt = - \beta n(A) m^{-3} s^{-1} \\ where \beta is photodissotiation rate of A

Chemical Kinetics

(**second order** - rate of formation and loss proportional to the concentration of two reactants)

 $\mathbf{A} + \mathbf{B} \rightarrow \mathbf{C} + \mathbf{D} \qquad \qquad \mathbf{k} = \langle \mathbf{\sigma} \mathbf{v} \rangle \quad \mathbf{m}^3 \ \mathbf{s}^{-1}$

 Loss of A and B per unit volume per unit second is:

dn(A)/dt = -kn(A)n(B) m⁻³ s⁻¹

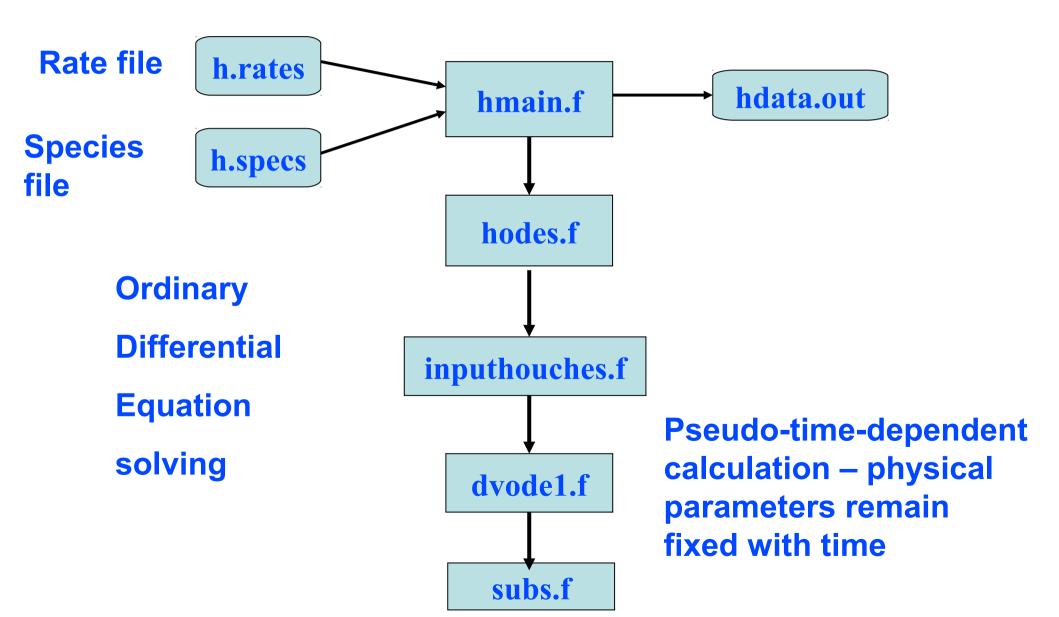
where n(A) is number of molecules of A per unit volume k and is rate constant

 Formation of C (and D) per unit volume per unit second is:

dn(C)/dt = + kn(A)n(B) m⁻³ s⁻¹

Kinetic Calculation

Millar *et. al.* 1991. Astron. Astrophys. Suppl. Ser. 87



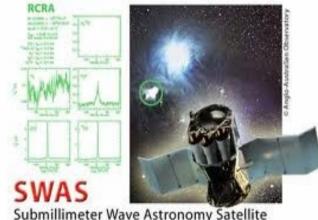
69 species



1	H+	0.00e+00	1.0								
2	H	0.00e+00	1.0								
з	H2+	0.00e+00	2.0	1	H	СН	с	H2	1.31E-10	0.00	80.0
4	H3+	0.00e+00	3.0	2	н	CH2	СН	H2 H2	6.64E-11	0.00	0.0
5	He+	0.00e+00	4.0	3	н	NH	N	H2	1.73E-11	0.50	2400.0
6	C+	0.00e+00	12.0	4	н	CH3	CH2	H2	1.00E-10	0.00	7600.0
7	c	0.00e+00	12.0	5	н	NH2	NH	H2	5.25E-12	0.79	2200.0
8	СН	0.00e+00	13.0	6	H	NH2	NH	H2	1.05E-10	0.00	4450.0
9	CH+	0.00e+00	13.0	7	H	CH4	CH3	H2	5.94E-13	3.00	4045.0
				8	н	OH	0	H2	6.99E-14	2.80	1950.0
10	CH2	0.00e+00	14.0	9	н	NH3	NH2	H2	7.80E-13	2.40	4990.0
11	CH2+	0.00e+00	14.0	10	H	H2O	OH	H2	1.59E-11	1.20	9610.0
12	N	0.00e+00	14.0	11	H	C2	CH	C	4.67E-10	0.50	30450.0
13	N+	0.00e+00	14.0	12	H	C2 H2	C2H	H2	3.80E-10	0.00	13634.0
14	NH	0.00e+00	15.0	13	H	HCN	CN	H2	6.20E-10	0.00	12500.0
15	CH3	0.00e+00	15.0	14 15	H H	С2 Н3 СО	C2 H2 OH	H2 C	3.32E-11	0.00	0.0 77700.0
16	CH3+	0.00e+00	15.0	16	н	нсо	o	C CH2	1.10E-10 6.61E-11	0.50 0.00	51598.0
17	NH+	0.00e+00	15.0	17	н	HCO	co	H2	2.00E-10	0.00	0.0
18	CH4+	0.00e+00	16.0	18	н	NO	0	NH	9.29E-10	-0.10	35220.0
19	NH2+	0.00e+00	16.0	19	н	NO	OH	N	3.60E-10	0.00	24910.0
				20	н	H2CO	HCO	H2	4.85E-12	1.90	1379.0
20	NH2	0.00e+00	16.0	21	н	HNO	OH	NH	2.40E-09	-0.50	9010.0
21	0+	0.00e+00	16.0	22	H	HNO	NH2	0	1.05E-09	-0.30	14730.0
22	CH4	0.00e+00	16.0	23	H	HNO	NO	H2	4.50E-11	0.72	329.0
23	0	0.00e+00	16.0	24	H	02	OH	0	2.61E-10	0.00	8156.0
24	OH+	0.00e+00	17.0	25	H2	С	CH	н	6.64E-10	0.00	11700.0
25	NH3+	0.00e+00	17.0	26	H2	CH	CH2	H	5.46E-10	0.00	1943.0
26	NH3	0.00e+00	17.0	27 28 29	H2	N	NH	H	1.69E-09	0.00	18095.0
27	OH	0.00e+00	17.0		H2	CH2	CH3	H	5.18E-11	0.17	6400.0
28	CH5+	0.00e+00	17.0		H2	NH	NH2	H	5.96E-11	0.00	7782.0
29	NH4+	0.00e+00	18.0								
30	H2O	0.00e+00	18.0								
	H2O+	0.00e+00	18.0								
32	H3O+	0.00e+00	19.0								

- Initial abundances: e, H₂, He, N, C, O, Mg
- Temperature, density, cosmic-rays...
- Gas-phase and grain surface chemistry
- Change some parameters to mach the observed abundances.
- SWAS satellite
 - $H_2^{0} \approx 6 \times 10^{-10} 1 \times 10^{-8}$

sensitive to C and O abundances



Subminieter wave Astronomy 5

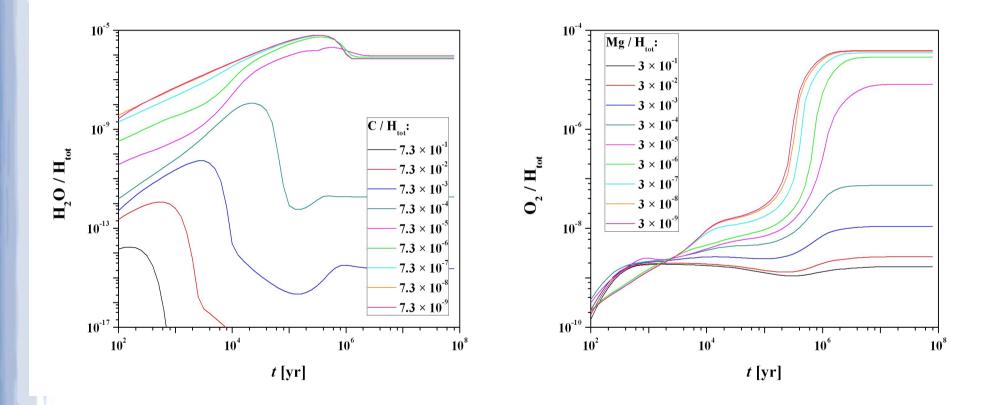
the best match for C $\approx 10^{\text{-5}}\text{--}10^{\text{-4}}$, O $\approx 10^{\text{-5}}$

•
$$O_2 \lesssim 10^{-7}$$

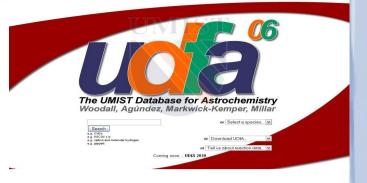
sensitive to C, O, N and Mg abundances

the best match for C $\approx 10^{\text{-5}}\text{--}10^{\text{-4}}$, O $\approx 10^{\text{-5}}\text{--}10^{\text{-4}}$, Mg $\approx 10^{\text{-4}}$, N $\approx 10^{\text{-2}}$

Initial abundances variation - example



Reaction rates



- Measured in laboratory or derived from theory
- On-line databases: UMIST database for Astrochemistry, Ohio State University (OSU), JPL Anicich Database, Huebner Photo-Cross-section Database...
- NIST Chemical Kinetics Database (VAMDC external partner)

27 000 reactions, rates from experiments and theory, generates best fit





Kinetics Database Resources

Simple Reaction Search

Search Reaction Database

Search Bibliographic Database

Set Unit Preferences

Feedback

National Institute of Standards and Technology

+ H2

Submit

Clear

NIST Chemical Kinetics Database

Standard Reference Database 17, Version 7.0 (Web Version), Release 1.6.5 Data Version 2012.02

A compilation of kinetics data on gas-phase reactions

Reaction Database Quick Search Form

Enter the reactant(s) and/or product(s) in the fields below. Fields may be left blank.

+ CH2

→ CH

If you would like more search options, try...

advanced reaction search form

bibliographic search form

H

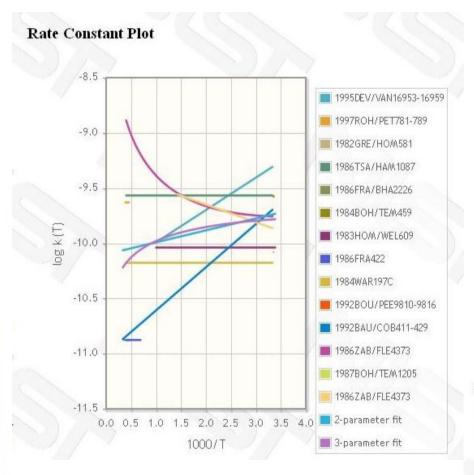
Search Results

Rate expression units: First order: s ⁻¹		8 K) ⁿ e ^{-Ea/RT}				
First order: s Second order: cm ³ /mole	cule s					
Third order: cm ⁶ /molecul R = 8.314472E-03 kJ / r			VIIa	10		
Energy Units	kJ	Molecular Units	Molecule			
Energy Units Pressure Units	kJ Pa	Molecular Units Temperature Units	Molecule K			

1

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	2
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.66E-10) 2
.00	2
1	

	1987BOH/TEM1205	$H \cdot + \cdot CH_2 \rightarrow H_2 + \cdot CH$	298	1.83E-10			1.83E-10	2	
	1986ZAB/FLE4373	$\begin{array}{c} H\cdot + \cdot CH_2 \rightarrow H_2 + \\ \cdot CH \end{array}$	300 - 2500	4.13E-11	1.54	-3.58		2	
	1986ZAB/FLE4373	$\begin{array}{l} \mathrm{H} \cdot + \cdot \mathrm{CH}_2 \longrightarrow \mathrm{H}_2 + \\ \cdot \mathrm{CH} \end{array}$	300 - 700	4.70E-10		3.08		2	
	1986FRA/BHA2226	$\begin{array}{c} \mathrm{H} \cdot + \cdot \mathrm{CH}_2 \longrightarrow \mathrm{H}_2 + \\ \cdot \mathrm{CH} \end{array}$	2000 - 2800	1.33E-11				2	
	1986FRA422	$H \cdot + \cdot CH_2 \rightarrow H_2 + \cdot CH$	1450 - 2500	1.33E-11				2	
	1984BOH/TEM459	$\begin{array}{c} \mathrm{H}\cdot + \cdot \mathrm{CH}_2 \longrightarrow \mathrm{H}_2 + \\ \cdot \mathrm{CH} \end{array}$	298	2.66E-10			2.66E-10	2	
	1983HOM/WEL609	$H \cdot + \cdot CH_2 \rightarrow H_2 + \cdot CH$	295 - 1000	9.13E-11			9.13E-11	2	
	1982GRE/HOM581	$\begin{array}{c} \mathrm{H} \cdot + \cdot \mathrm{CH}_2 \longrightarrow \mathrm{H}_2 + \\ \cdot \mathrm{CH} \end{array}$	298	8.30E-11			8.30E-11	2	
Theory									
	1987BEC/CAN435-448	$\begin{array}{c} \mathrm{H} \cdot + \cdot \mathrm{CH}_2 \rightarrow \mathrm{H}_2 + \\ \cdot \mathrm{CH} \end{array}$							



Fit of Arrhenius parameters to set:

Temperature range: 295 - 3000 K

Two-parameter fit:

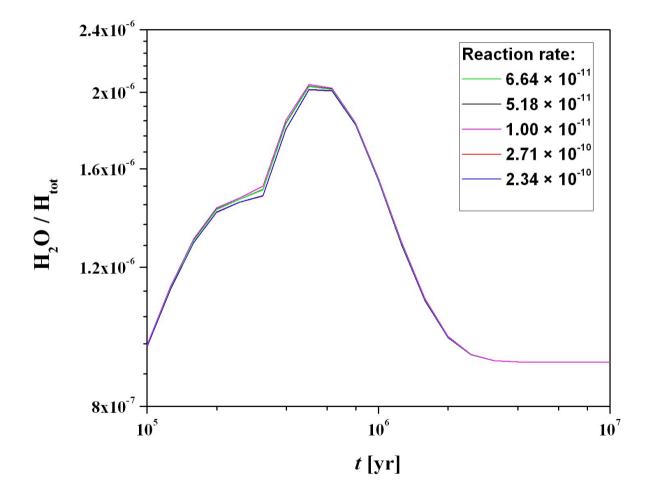
 $k(T) = Aexp(-E_a/RT)$

 $k(T) = A(T/T_{\rm ref})^n exp(-E_a/RT)$

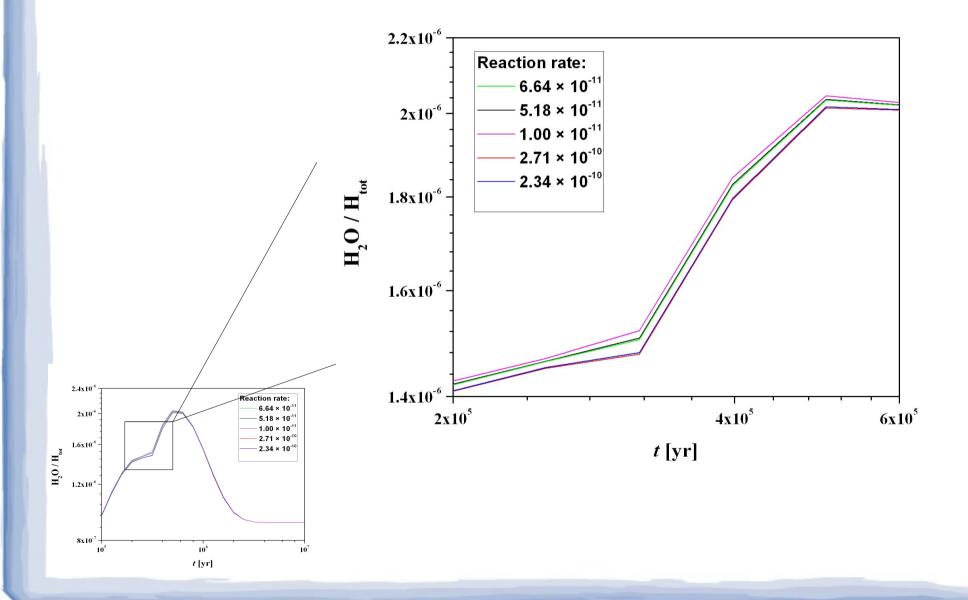
 $A = 7.94 \times 10^{-11}$ [cm³/molecule s] $E_a = -2.1$ [kJ] RMSD = 1.4 $A = 2.26 \times 10^{-10} \text{ [cm^3/molecule s]}$ n = -0.6 $T_{ref} = 298 \text{ [K]}$ $E_a = 0.8 \text{ [kJ]}$ RMSD = 1.3

Three-parameter fit:

Reaction rates - example



Reaction rates - example



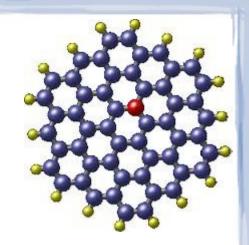
Conclusion



- Chemical databases very important for astrochemistry
- Modeling of chemical evolution of ISM, for example dark clouds
- Many reaction to be considered
- Data needed for calculations can be found in on-line databases
- Faster and easier search is important







Thank you for your attention!

