

# VAMDC as a Resource for Atomic and Molecular Data and the New Release of VALD

**Friedrich Kupka**

Faculty of Mathematics, University of Vienna, Austria

and the **VAMDC Collaboration (P.I. M.-L. Dubernet)**

M.L. Dubernet et al. 2010, JQSRT 111, 2152



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**T. Ryabchikova**

INASAN, Moscow, Russia

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**N. Piskunov, U. Heiter, H.C. Stempels**

from the VALD team in Uppsala, Sweden, for providing various figures.

Further input has been taken from the VALD and VAMDC homepages at <http://vald.astro.univie.ac.at/~vald/php/vald.php> and <http://www.vamdc.eu/> as well as from presentations by **N. Mason** and **G. Rixon** at the 2<sup>nd</sup> annual VAMDC Conference.

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# OUTLINE

- Part 1: VAMDC
  - Node members
  - What is it ? Idea behind it ? Expected outcome ?
  - Clientele, astrophysical examples
  - Technology, available databases
- Part 2: VALD
  - The Team
  - Current status, concepts and facilities, structure
  - VALD-3 data
  - VALD-3 in action and shortcomings

# VAMDC Vienna Node Team Members

**Friedrich Kupka** → **Institute of Mathematics**

**Theresa Rank-Lüftinger** → **Institute for Astronomy**

**Werner W. Weiss** → **Institute for Astronomy**

and until May 2011

**Christian Stütz** → **Computer Centre of the University**

# What is VAMDC ?

The Virtual Atomic and Molecular Data Centre (VAMDC) is an international project aimed to create an interoperable interface to existing Atomic and Molecular (A&M) databases. In general, each database within VAMDC has been organised in its own way.

It consists of 8 work packages. 15 administrative partners which represent 24 teams from 6 countries within the EU (Austria, France, Germany, Italy, Sweden, and the UK) as well as from Russia, Serbia, and Venezuela have entered the VAMDC consortium.

Presently, the VAMDC project includes 24 A&M DBs, among others: **BASECOL, CDMS, CDSD, CHIANTI, Ethelyne, GhoSST, HITRAN, KIDA, PAH, SPECTR-W<sup>3</sup>, StarkB, TIPbase, TOPbase, VALD.**

A detailed description of VAMDC is given by ***Dubernet et al. (2010), JQSRT 111, 2151.***

# The Idea behind VAMDC

- A&M data have been collected and assessed in a various databases. They underpin a wide range of physics in applied research and industrial development.
- Many databases have been built to serve specific needs. VALD is one such example with its own advantages, special tools, and limitations.
- Various kinds of data, formats, completeness, etc. and
- specialized extraction tools exist for each of the A&M databases.
- Problems:
  - data duplication
  - different user interfaces
  - restricted access
  - often fragmentary, ...

# Expected VAMDC outcomes

- Develop/extend standards for interoperability of A&M resources
- Implementation of selected databases
- Find resources easily
- Query those resources with dedicated protocols/languages
- Transfer large amounts of data asynchronously
- Create a safe environment for publishing latest sets of A&M data
- Linking producers with users
  
- Key benefits from using VAMDC:
  - find any type of AM data with a click, provide uniform access
  - cross-matching different data sets, wide access to latest published data

# A & M Users: The VAMDC Clientele

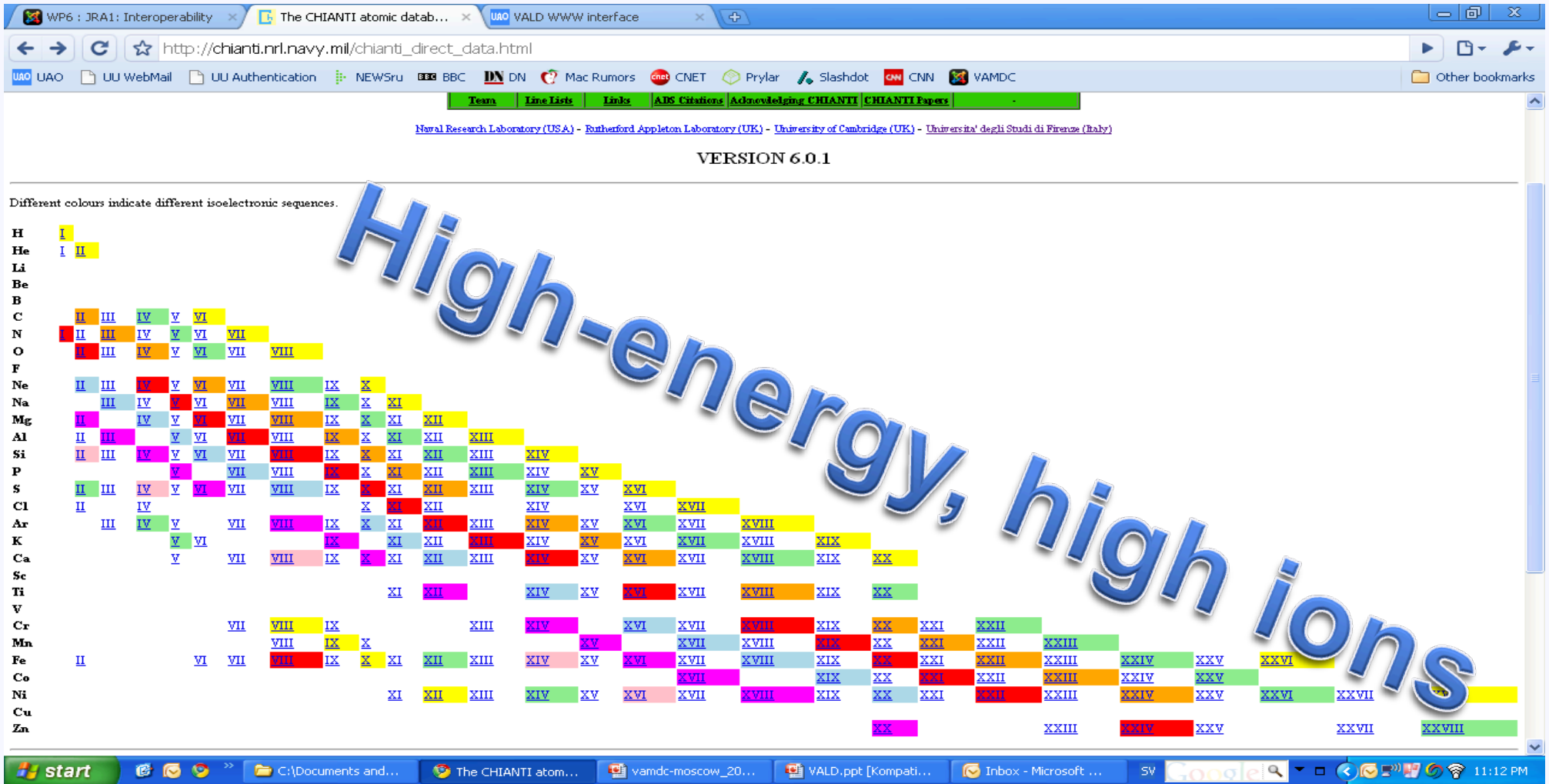
- Astrophysics / Astronomy / Planetary Science
- Atmospheric Science
- Fusion Science
- Plasma Science
- Radiation Science

and their applications in research and industrial development



# Astronomical Examples I

## Chianti atomic line database (example courtesy of N. Piskunov)



The screenshot shows the Chianti atomic line database website. The browser address bar displays `http://chianti.nrl.navy.mil/chianti_direct_data.html`. The page title is "VERSION 6.0.1". A navigation menu includes "Team", "Line Lists", "Links", "ADS Citations", "Acknowledging CHIANTI", and "CHIANTI Papers". A list of contributing institutions is provided: "Naval Research Laboratory (USA) - Rutherford Appleton Laboratory (UK) - University of Cambridge (UK) - Universita' degli Studi di Firenze (Italy)".

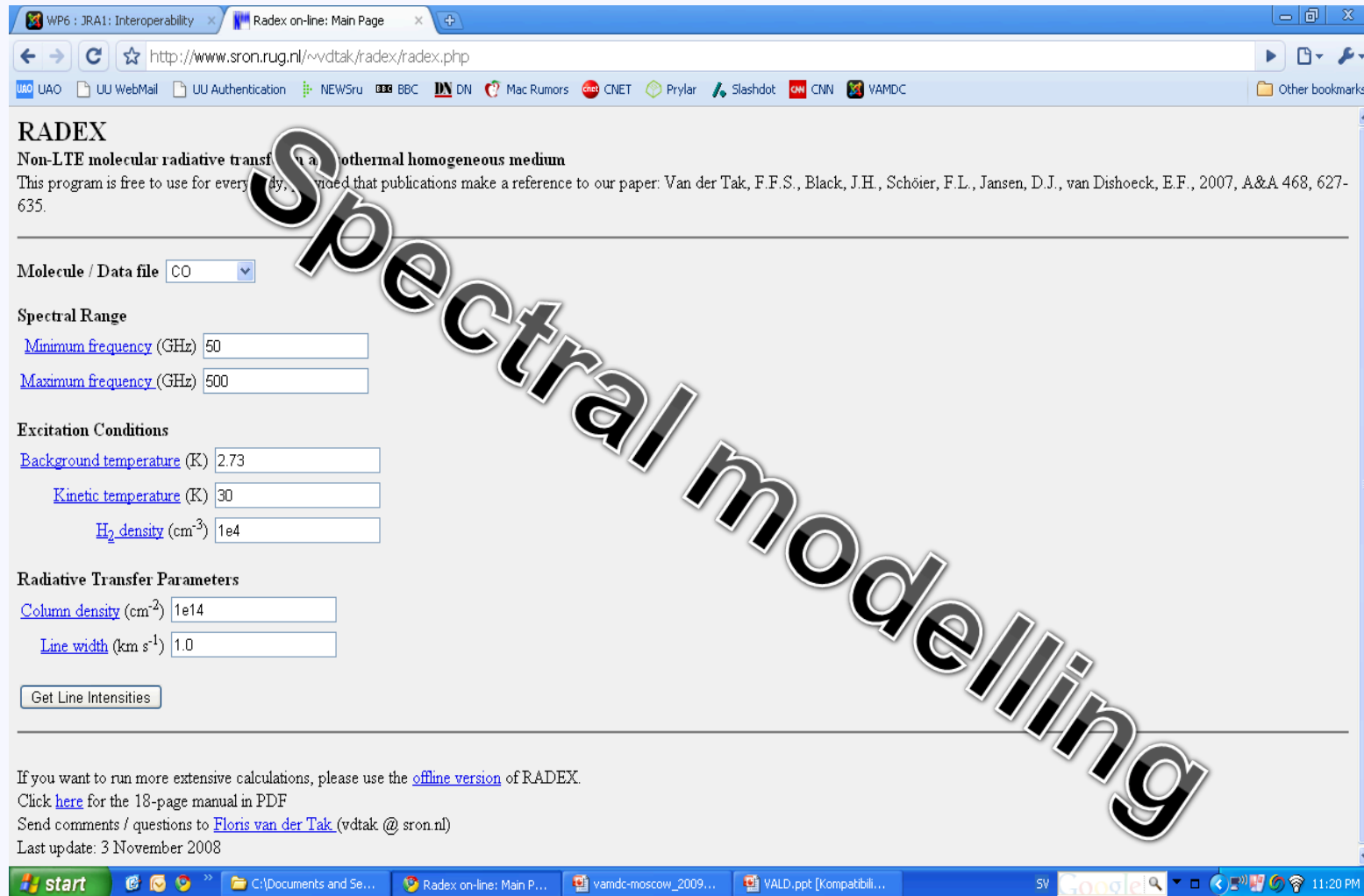
The main content is a chart showing ionization sequences for elements from Hydrogen (H) to Zinc (Zn). The chart is a grid where rows represent elements and columns represent ionization states (I, II, III, etc.). Different colors indicate different isoelectronic sequences. A large blue watermark "High-energy, high ions" is overlaid diagonally across the chart.

Legend: Different colours indicate different isoelectronic sequences.

| Element | I | II | III | IV  | V  | VI  | VII  | VIII | IX   | X  | XI  | XII  | XIII | XIV  | XV  | XVI  | XVII  | XVIII | XIX | XX  | XXI  | XXII  | XXIII | XXIV | XXV  | XXVI | XXVII | XXVIII |  |
|---------|---|----|-----|-----|----|-----|------|------|------|----|-----|------|------|------|-----|------|-------|-------|-----|-----|------|-------|-------|------|------|------|-------|--------|--|
| H       | I |    |     |     |    |     |      |      |      |    |     |      |      |      |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| He      | I | II |     |     |    |     |      |      |      |    |     |      |      |      |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| Li      |   |    |     |     |    |     |      |      |      |    |     |      |      |      |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| Be      |   |    |     |     |    |     |      |      |      |    |     |      |      |      |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| B       |   |    |     |     |    |     |      |      |      |    |     |      |      |      |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| C       |   | II | III | IV  | V  | VI  |      |      |      |    |     |      |      |      |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| N       |   | II | III | IV  | V  | VI  | VII  |      |      |    |     |      |      |      |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| O       |   | II | III | IV  | V  | VI  | VII  | VIII |      |    |     |      |      |      |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| F       |   |    |     |     |    |     |      |      |      |    |     |      |      |      |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| Ne      |   | II | III | IV  | V  | VI  | VII  | VIII | IX   | X  |     |      |      |      |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| Na      |   |    | III | IV  | V  | VI  | VII  | VIII | IX   | X  | XI  |      |      |      |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| Mg      |   |    | II  | III | IV | V   | VI   | VII  | VIII | IX | X   | XI   | XII  |      |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| Al      |   |    | II  | III | IV | V   | VI   | VII  | VIII | IX | X   | XI   | XII  | XIII |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| Si      |   |    | II  | III | IV | V   | VI   | VII  | VIII | IX | X   | XI   | XII  | XIII | XIV |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| P       |   |    |     | III | IV | V   | VI   | VII  | VIII | IX | X   | XI   | XII  | XIII | XIV | XV   |       |       |     |     |      |       |       |      |      |      |       |        |  |
| S       |   |    | II  | III | IV | V   | VI   | VII  | VIII | IX | X   | XI   | XII  | XIII | XIV | XV   | XVI   |       |     |     |      |       |       |      |      |      |       |        |  |
| Cl      |   |    | II  | III | IV | V   | VI   | VII  | VIII | IX | X   | XI   | XII  | XIII | XIV | XV   | XVI   | XVII  |     |     |      |       |       |      |      |      |       |        |  |
| Ar      |   |    | III | IV  | V  | VI  | VII  | VIII | IX   | X  | XI  | XII  | XIII | XIV  | XV  | XVI  | XVII  | XVIII |     |     |      |       |       |      |      |      |       |        |  |
| K       |   |    |     |     | V  | VI  | VII  | VIII | IX   | X  | XI  | XII  | XIII | XIV  | XV  | XVI  | XVII  | XVIII | XIX |     |      |       |       |      |      |      |       |        |  |
| Ca      |   |    |     |     | V  | VI  | VII  | VIII | IX   | X  | XI  | XII  | XIII | XIV  | XV  | XVI  | XVII  | XVIII | XIX | XX  |      |       |       |      |      |      |       |        |  |
| Sc      |   |    |     |     |    |     |      |      |      |    |     |      |      |      |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| Ti      |   |    |     |     |    |     |      |      |      |    | XI  | XII  | XIII | XIV  | XV  | XVI  | XVII  | XVIII | XIX | XX  |      |       |       |      |      |      |       |        |  |
| V       |   |    |     |     |    |     |      |      |      |    |     |      |      |      |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| Cr      |   |    |     |     |    | VII | VIII | IX   | X    | XI | XII | XIII | XIV  | XV   | XVI | XVII | XVIII | XIX   | XX  | XXI | XXII |       |       |      |      |      |       |        |  |
| Mn      |   |    |     |     |    |     | VIII | IX   | X    | XI | XII | XIII | XIV  | XV   | XVI | XVII | XVIII | XIX   | XX  | XXI | XXII | XXIII |       |      |      |      |       |        |  |
| Fe      |   | II |     |     | VI | VII | VIII | IX   | X    | XI | XII | XIII | XIV  | XV   | XVI | XVII | XVIII | XIX   | XX  | XXI | XXII | XXIII | XXIV  | XXV  | XXVI |      |       |        |  |
| Co      |   |    |     |     |    |     |      |      |      |    |     |      |      |      |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| Ni      |   |    |     |     |    |     |      |      |      |    |     |      |      |      |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| Cu      |   |    |     |     |    |     |      |      |      |    |     |      |      |      |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |
| Zn      |   |    |     |     |    |     |      |      |      |    |     |      |      |      |     |      |       |       |     |     |      |       |       |      |      |      |       |        |  |

# Astronomical Examples II

## Leiden Atomic and Molecular Database (example courtesy of N. Piskunov)



**RADEX**  
Non-LTE molecular radiative transfer in a thermal homogeneous medium  
This program is free to use for everybody, provided that publications make a reference to our paper: Van der Tak, F.F.S., Black, J.H., Schöier, F.L., Jansen, D.J., van Dishoeck, E.F., 2007, A&A 468, 627-635.

Molecule / Data file:

**Spectral Range**  
 Minimum frequency (GHz):   
 Maximum frequency (GHz):

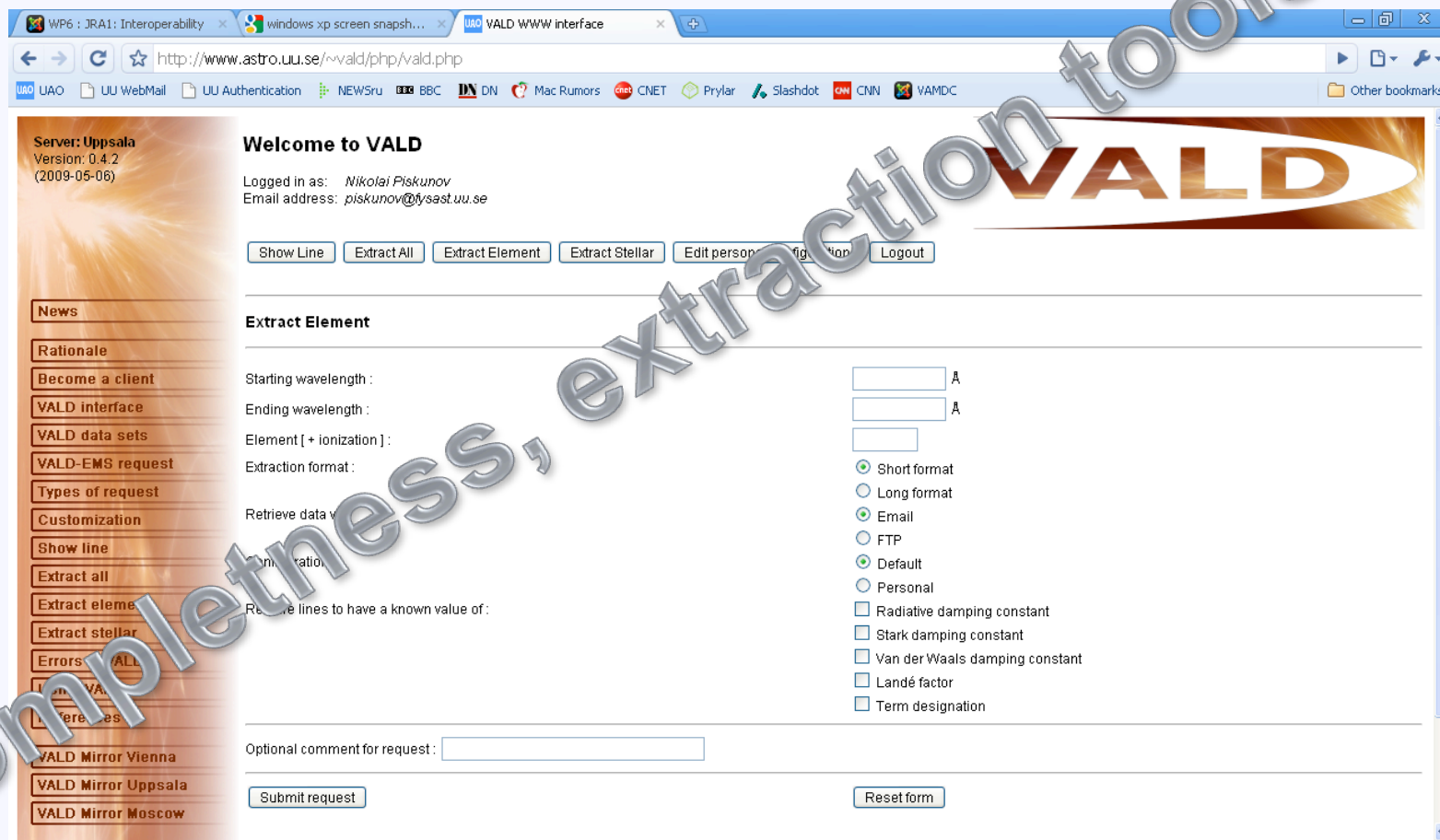
**Excitation Conditions**  
 Background temperature (K):   
 Kinetic temperature (K):   
 H<sub>2</sub> density (cm<sup>-3</sup>):

**Radiative Transfer Parameters**  
 Column density (cm<sup>-2</sup>):   
 Line width (km s<sup>-1</sup>):

If you want to run more extensive calculations, please use the [offline version](#) of RADEX.  
 Click [here](#) for the 18-page manual in PDF  
 Send comments / questions to [Floris van der Tak](mailto:vdtak@sron.nl) (vdtak@sron.nl)  
 Last update: 3 November 2008

# Astronomical Examples III

Vienna  
Atomic  
Line  
Database



Server: Uppsala  
Version: 0.4.2  
(2009-05-06)

**Welcome to VALD**  
Logged in as: *Nikolai Piskunov*  
Email address: *piskunov@fysast.uu.se*

Show Line Extract All Extract Element Extract Stellar Edit personal configuration Logout

**Extract Element**

Starting wavelength :  Å

Ending wavelength :  Å

Element [ + ionization ] :

Extraction format :  Short format  
 Long format  
 Email  
 FTP  
 Default  
 Personal

Radiative damping constant  
 Stark damping constant  
 Van der Waals damping constant  
 Landé factor  
 Term designation

Optional comment for request :

Submit request Reset form

(example courtesy of N. Piskunov)

# Technology I: VAMDC Components

“Grand Central” in Paris

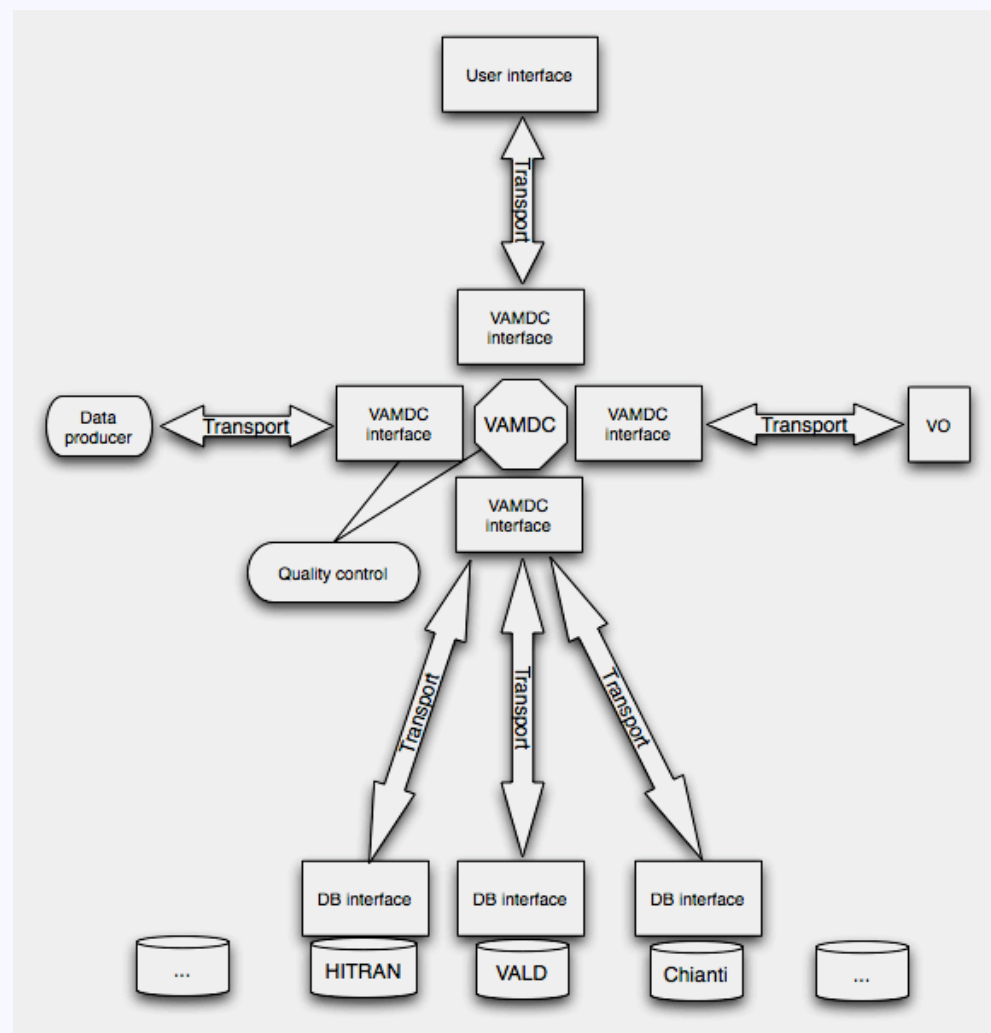
Web-based transport protocol

Interface(s) to DBs

User interface

Automatic interface (e.g. to VO)

Data publishing tools



# Technology II: Web-Based Transport Protocol

- The transport protocol handles:
  - database (DB) queries for status and data content (registry update)
  - query / data transfer between user and Grand Central
  - query / data transfer between a DB and Grand Central
- The transport protocol is:
  - self-descriptive (XML description of what is sent including units, formats etc., use of XSAMS data format)
  - efficient for large data sets (compressed binary tables)

# Technology III: Interface to Databases

- Fully compatible with the VAMDC transport on the outside
- Tuned to the specific DB on the inside:
  - converting incoming queries to the internal query format
  - converting the DB extraction to the transport-compatible format
- Capable of responding to VAMDC-specific queries (registry update)
- Capable of collecting the accounting information

# Technology IV: Registry Browser



The screenshot shows the VAMDC Registry Browser interface. At the top left is the VAMDC logo. Below it is a navigation bar with links: VAMDC Portal, Home, XSAMS Query Builder, and Query Log. The main content area is divided into two sections: 'Primary Options' and 'Atoms'. The 'Primary Options' section has three sub-sections: 'Species', 'Wavelength', and 'Free Form'. The 'Atoms' section contains three input fields: 'Atomic (elemental) symbol:', 'Atom Ion Charge: from [ ] to [ ]', and 'Atom Mass Number: from [ ] to [ ]'. Below these is a 'Wavelength' section with a 'RadTransWavelength: from [ ] to [ ]' input field. At the bottom of the form are three buttons: 'Clear', 'Cancel', and 'Find suitable databases'. Below the form is a footer with the text 'Developed in IoA, University of Cambridge by Asif Akram.' and a row of logos from various partner institutions including the European Union, UCL, University of Cambridge, and others.

snapshot of webpage of test version as of 3 June 2011

# Databases for Cycle 2 Service Release I

| Database            | Node     | Included? |
|---------------------|----------|-----------|
| KIDA                | Bordeaux | Yes       |
| Methane lines       | Dijon    | Yes       |
| VALD                | Uppsala  | Yes       |
| BASECOL             | LPMAA    | Yes       |
| S&MPO               | Reims    | Yes       |
| Ethelyne            | Reims    | Yes       |
| GhoSST              | Grenoble | Yes       |
| CDSD                | IAO      | Yes       |
| eMOL                | OU       | Maybe     |
| Line lists          | UCL      | No        |
| Energy levels       | UCL      | No        |
| Electron collisions | UCL      | No        |

Status according to Guy Rixon at 2<sup>nd</sup> VAMDC annual conference in March 2011



# Databases for Cycle 2 Service Release II

| Database              | Node           | Included?      |
|-----------------------|----------------|----------------|
| PAH                   | OAC            | Probably       |
| Mols./Mixtures/Solids | OAC            | Probably       |
| UDMA                  | QUB            | Maybe          |
| CDMS                  | Koln           | Yes            |
| Stark-B               | OPM            | Yes            |
| Stark-B               | AOB            | Maybe (mirror) |
| Stark-C               | OPM            | No             |
| TipTopBase            | IVIC/Cambridge | Yes            |
| uADB                  | IVIC/Cambridge | Probably       |
| CHIANTI               | Cambridge/MSSL | Yes            |
| HITRAN                | UCL            | Yes            |
| Spectr-W              | RFNC-VNIITF    | Yes            |
| Lund data             | Uppsala        | Yes            |

Status according to Guy Rixon at 2<sup>nd</sup> VAMDC annual conference in March 2011

- **VALD Team and Server Locations**

- **Vienna:** Inst. for Astronomy ([founding site](#), server main site)
- **Uppsala:** Observatory (mirror site, [now main developer site](#))
- **Moscow:** Institute of Astronomy RAS (mirror site, [now main data collection site](#))
- non-public mirror sites: GSFC, STSI Baltimore, AIP Potsdam (VALD-2, US sites no longer maintained for practical reasons)

- **Team Members of VALD-2**

P. Barklem, F. Kupka, N.E. Piskunov, T.A. Ryabchikova,  
H.C. Stempels, W.W. Weiss

- **Collaborators & Data providers of VALD-2**

CCP7, NIST, R.L. Kurucz, C.R. Cowley (initially, followed by many more in the mean time → VALD-3)

- **Team Members of VALD-3 (Software, Core Team)**
  - **Vienna:** F. Kupka, T. Rank-Lüftinger, W.W. Weiss  
(previous members: N. Nesvacil, M. Obbrugger, Ch. Stütz)
  - **Uppsala:** U. Heiter, N. Piskunov, H.C. Stempels  
as well as P. Barklem and O. Kochukhov (+ VAMDC team)
  - **Moscow:** T. Ryabchikova, Yu. Pakhomov (+ data providers)
- **Members of the VALD-3 data providing teams**
  - **Univ. of Wisconsin:** J.E. Lawler, E.A. Den Hartog, et al. (REE)
  - **Lund University:** Lund team (Fe peak, Th+U)
  - **DREAM Database:** E. Biemont, et al. (REE)
  - **Univ. of Montpellier:** B. Plez (molecules)

# VALD Concepts and Facilities I

## Main publications on VALD-1 and VALD-2

- VALD-1: Piskunov N.E., Kupka F., Ryabchikova T.A., Weiss W.W., Jeffery C.S., [A&AS 112, 525 \(1995\)](#)
- VALD-2: Kupka F., Piskunov N.E., Ryabchikova T.A., Stempels H.C., Weiss W.W., [A&AS 138, 119 \(1999\)](#)
- VALD-2/Overview: Ryabchikova T.A., Piskunov N.E., Stempels H.C., Kupka F., Weiss W.W., [Physica Scripta T83, 162 \(1999\)](#)
- **VALD-3: in preparation**

# VALD Concepts and Facilities II

## Goal of the VALD project

- compile **accurate and complete** line lists
  - for stellar atmospheres & spectroscopy
- evaluate line lists → **provide a ranking**
- provide a database which features
  - expandability with respect to data and contents
  - simple access through “customized” extraction software
  - fast access to individual data entries
  - an overview of parameters from different sources
  - compilation of data references and provision of quality criteria
  - to extract sets of best data according to data ranking lists

# VALD Concepts and Facilities III

## General architecture

- **Standard data format**
  - units which are common in astrophysics
  - one record with fixed length per spectral line
- **Semi-direct access** (compressed binary line lists)
- **Ranking** (for merging data)
- **Multiple extraction layers within VALD-2:**
  - access files / merge VALD data
  - prepare output for applications
  - remote access (EMS)
  - Web interface for EMS

# VALD Concepts and Facilities IV

## File format of VALD-2 data

- each line of each list: 1 record, fixed length
- junks of 1024 lines: compressed, index file ( $\lambda$  sorted), compression factor  $\sim 25\dots 40$

## One VALD-2 Data record is one VALD line

- **Mandatory entries (no defaults)**
  - central  $\lambda$ , species identifier,  $\log(gf)$ ,  $E_i$ ,  $J_i$ ,  $E_k$ ,  $J_k$
- **Optional entries (defaults exist)**
  - $g_i$ ,  $g_k$ ,  $\log(\Gamma_r)$ ,  $\log(\Gamma_s)$ ,  $\log(\Gamma_w)$ , terms (i,k), accuracy, comments (multiplets, e.g.)
  - flags (links to specific data for a line & to other data bases)

# VALD Concepts and Facilities V

## Ranking Lists

- generic extraction constraints
  - wavelength window: species + J values +  $\Delta E_k < 0.1\%$   $\rightarrow \Delta \lambda$
  - max. spectrum number
  - max. excitation potential
- extraction constraints for merging line data
  - VALD internal index number (switch list status to on / off)
  - element range of list
  - ranking for  $\lambda$ ,  $gf$ ,  $E_i$ ,  $E_k$ ,  $\langle g_{\text{eff}} \rangle$ ,  $\Gamma_r$ ,  $\Gamma_s$ ,  $\Gamma_w$ , level classification
- VALD **default** ranking or **user defined** ranking



# VALD Concepts and Facilities VI

## Extraction constraints & ranking

- Line found in different source lists ?
  - choose for each atomic parameter the value from the list with the highest ranking
- Line lists
  - homogeneous, high quality → high ranking
  - inhomogeneous or low quality → low ranking
  - VALD made lists for homogenization → usually high ranking
  - separate line lists usually correspond to separate data files

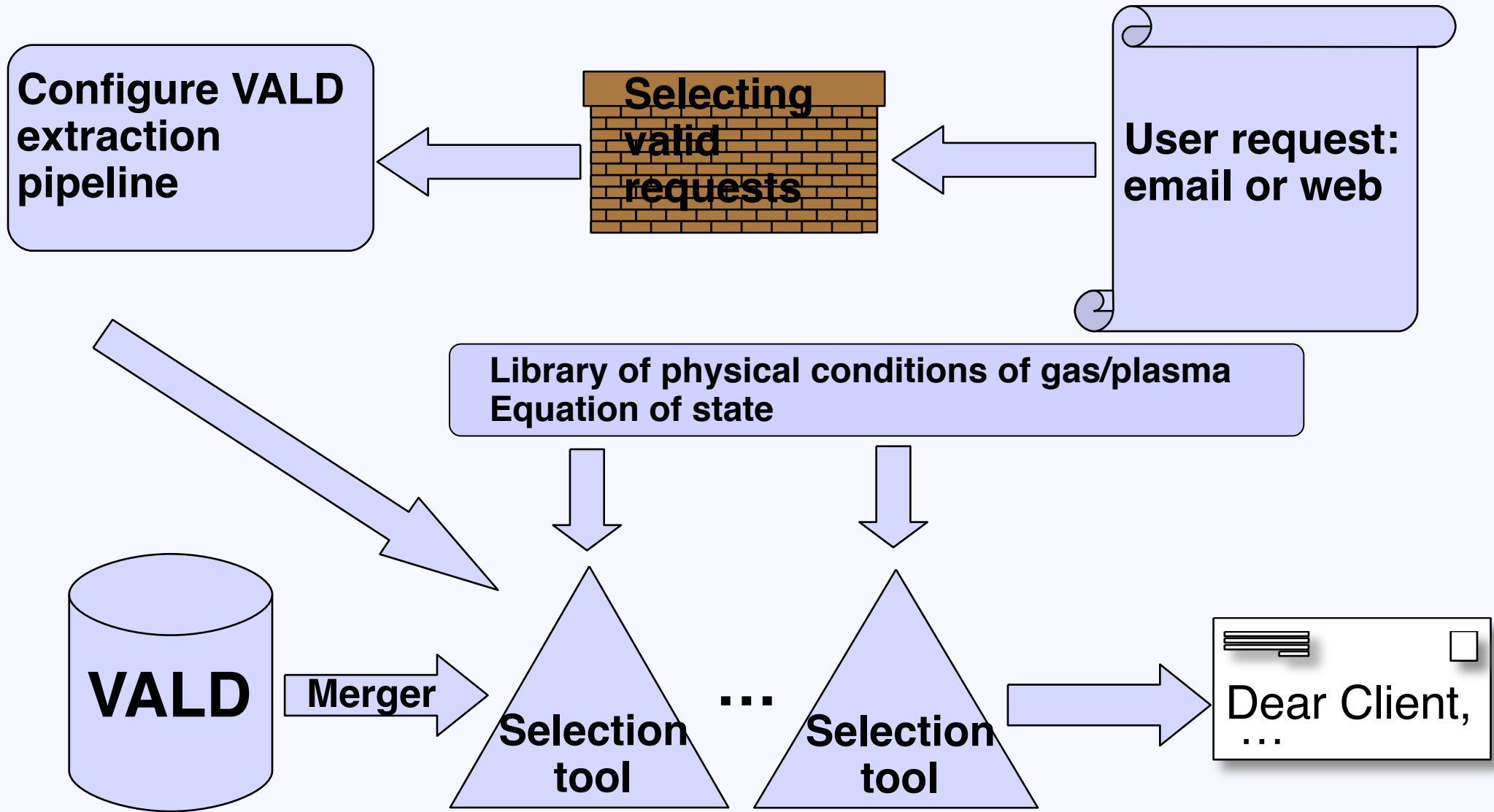
# VALD Concepts and Facilities VII

## Quality determination based on

- Error estimates from original sources
- Intercomparison of existing, alternative sources
- Applications in astrophysics (user feedback)
- Reranking possible
- General guidelines
  - prefer **experimental** data over calculations  
(with few exceptions such as the work of Uyling & Raassen)
  - prefer data with **individual error estimates**

## VALD Structure

### VALD from outside...



# In Preparation of VALD-3: The Current Status of VALD

- Over 150 line lists, over 66 million atomic lines provided by all major spectroscopy centers across the world
- Mirror sites in Vienna, Uppsala and Moscow
- Close to 1500 users from more than 50 countries, in average 30 requests processed per day

## VALD-3 – The new VALD, Part I

- data is still sorted as a function of wavelength and
- still stored in a special compressed format with semi-direct access
- stored data (changes & additions indicated) contains:
  - species, wavelengths ( $\text{\AA}$ , **vac**), level energies (now  $\text{cm}^{-1}$ ), total angular momentum quantum number,  $\log gf$ , Landé-factors, damping constants, **accuracy in  $\log gf$** , data reference, **full level designation and term name**
- publishing a new dataset in VALD means adding a new data file
- data description stored in various support files (list of species, configuration file)
- configuration file stores ranks for every field in each file
- reference data of each data set provided in **BibTeX format**

## VALD-3 – The new VALD, Part II

- in addition to many smaller new lists from data providers previously mentioned VALD-3 contains the
- New Kurucz Calculations (2006-2010) for Fe-peak elements
- Model-based selection improved:
  - for a set P-T-[abundance] VALD-3 solves the equation of state and estimates the contribution to opacities,
  - if a sequence of P-T is available (e.g. model atmosphere), VALD-3 will solve the radiative transfer to predict the line strength.

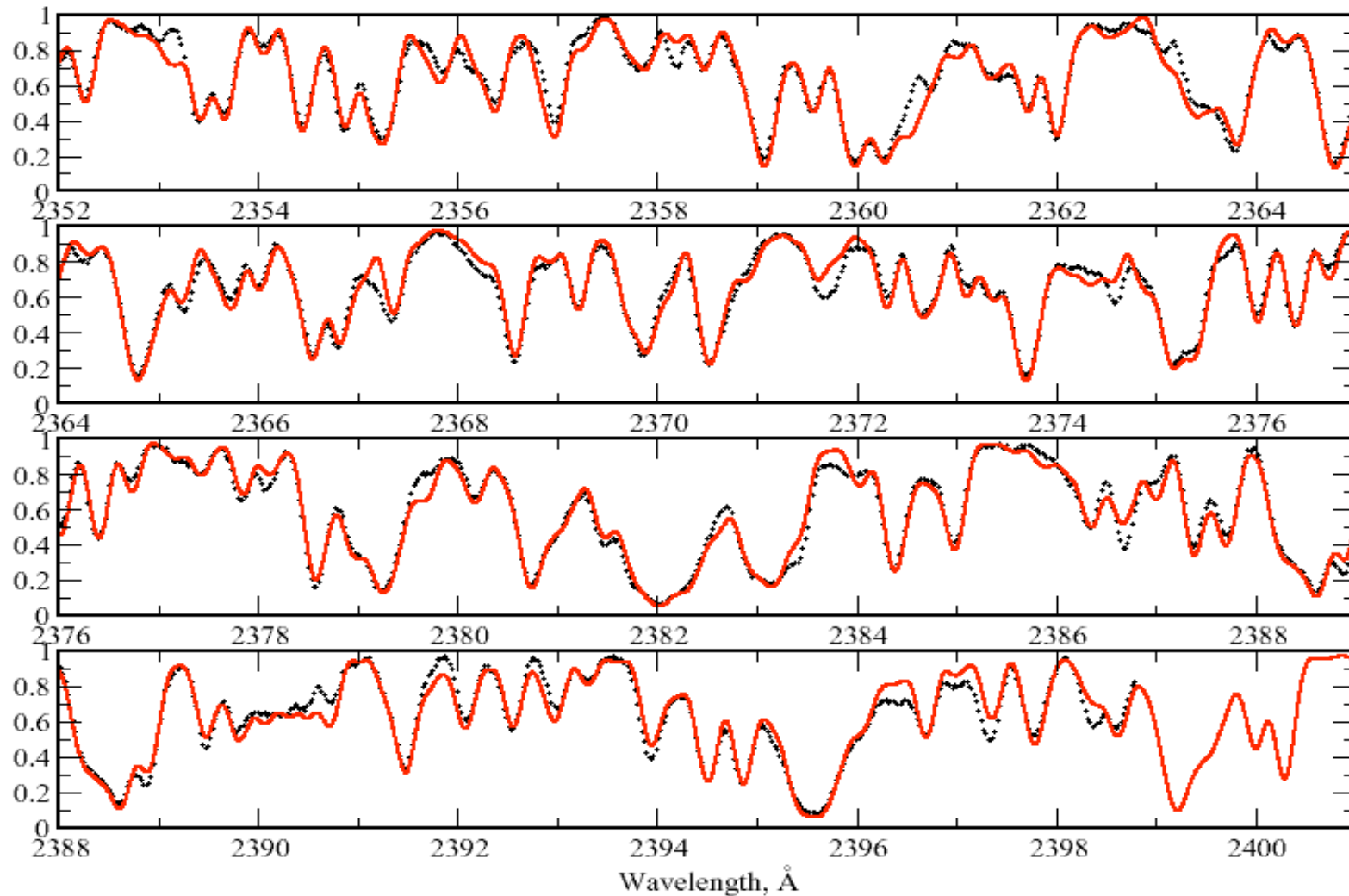
# VALD-3 – The new VALD, Part III

Number of lines in new (experimental and predicted) and old (total) Kurucz data:

|        | new   |          | old     |        | new    |           | old            |
|--------|-------|----------|---------|--------|--------|-----------|----------------|
|        | exp   | pred     |         |        | exp    | pred      |                |
| Sc I   | 15546 | 737992   | 191253  | Fe I   | 93508  | 6029023   | 789176         |
| Sc II  | 3436  | 116491   | 49811   | Fe II  | 103357 | 7615097   | 1264969        |
| Sc V   | 2180  | 645368   | 130563  | Fe III | 37199  | 9548787   | 1604934        |
| Ti I   | 33815 | 4758992  | 867399  | Co I   | 15441  | 3771900   | 546130         |
| Ti II  | 8188  | 835027   | 264867  | Co II  | 23355  | 10050728  | 1361114        |
| Ti III | 4090  | 499739   | 23742   | Co III | 9356   | 11515139  | 2198940        |
| V I    | 23342 | 7043556  | 1156790 | Ni I   | 9663   | 732160    | 149925         |
| V II   | 18389 | 3932853  | 925330  | Ni II  | 55590  | 3645991   | 404556         |
| V III  | 9892  | 966528   | 284003  | Ni III | 21251  | 11120833  | 1309729        |
| Cr I   | 35315 | 2582957  | 434773  | Total  | 623360 | 115185086 | 21778816 5.3/1 |
| Cr II  | 58996 | 6970052  | 1304043 |        |        |           |                |
| Cr III | 23150 | 5535931  | 990951  |        |        |           |                |
| Mn I   | 16798 | 1481464  | 327741  |        |        |           |                |
| Mn II  | 31437 | 4523390  | 878996  |        |        |           |                |
| Mn III | 17294 | 10525088 | 1589314 |        |        |           |                |

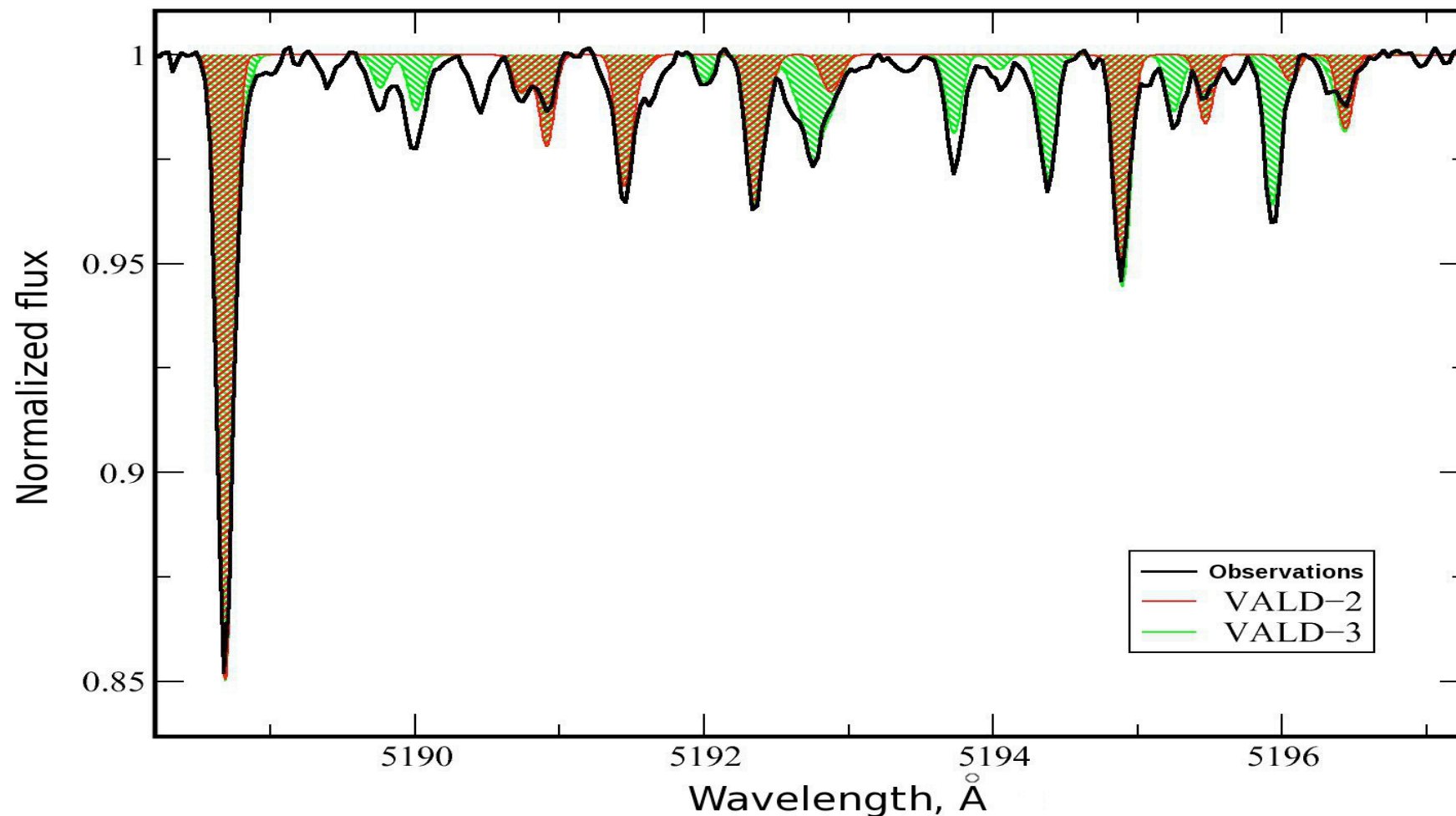
(example by courtesy of T. Ryabchikova)

# VALD-3 in Action I



Sirius A ([GHR@HST](#)): fit to UV region observations (courtesy John Landstreet), example courtesy of T. Ryabchikova

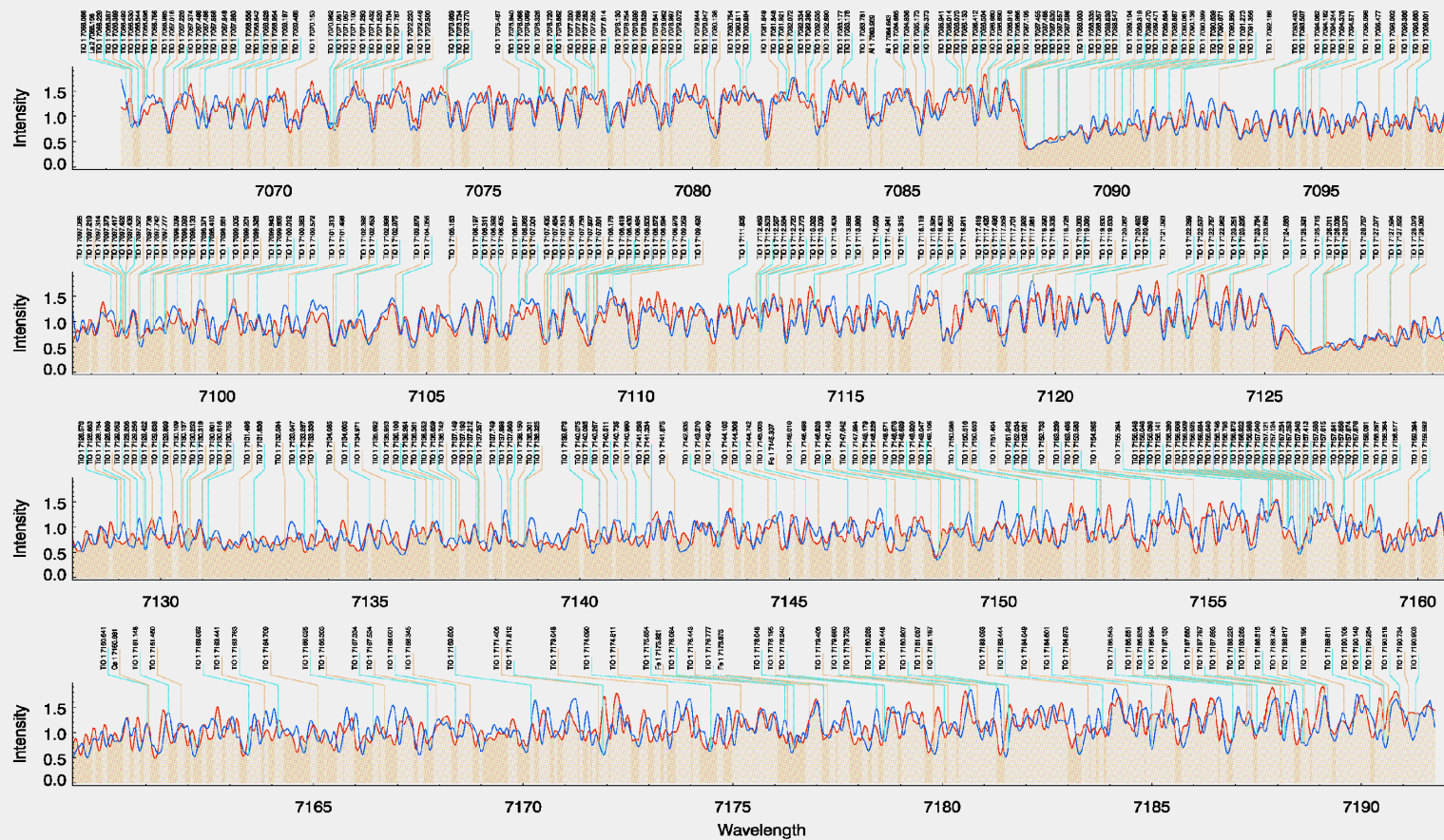




Solar composition star 21 Peg ( $T_{\text{eff}}=10400$  K): fit to optical region observations, example courtesy of T. Ryabchikova

# VALD-3 in Action III

gj876 metallicity fit



Molecular lines in the near infrared spectrum of an M dwarf (low mass star), (example courtesy of N. Piskunov & the VALD team)

# VALD-3 Shortcomings

- Limited range of ionization stages (neutral up to 8 times).
- Only simple molecules will be included in VALD-3 (basically diatomics: TiO, CO, CN, CH, FeH).
- Generally missing data include collisional transition probabilities, advanced broadening approximations.
- VALD consortium has no manpower or expertise to fix these deficiencies !

→ possible solution: access VAMDC data resources !

# Upcoming Conferences which will provide training on using VAMDC

3<sup>rd</sup> annual VAMDC Conference, Vienna, Austria, 21-24 February 2012

(see <http://www.vamdc.eu/>)

VAMDC Regional Workshop and School in Atomic and Molecular Data, Belgrade, Serbia, 7-9 June 2012

(see <http://poincare.matf.bg.ac.rs/~andjelka/VAMDC/>)

**...THANK YOU FOR YOUR TIME !**