XIV SCSLSA OVERVIEW OF LARGE SPECTROSCOPIC SURVEYS

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THE ERA OF LARGE SURVEYS

DEEP-WIDE field imaging surveys coming online in the next years

Vera Rubin LSST (NOIRLAb): 6.5m, 10 years, ugrizy ~ 25-27 over 18,000 deg²

Nancy Roman Space Telescope (NASA / ex WFIRST): 2.4m, YJH ~ 26-27 over 2,000 deg²

Euclid Space Telescope (ESA): 1.2m, 6 year mission, riz+YJH ~ 24 over 15,000 deg²





ugrizy ~ 25-27 over 18,000 deg² ugrizy ~ 27-29 over 38 deg² (deep drilling fields)

ission, riz+YJH ~ 24 over 15,000 deg² ~ 26 over 40 deg²







THE ERA OF LARGE SURVEYS

DEEP-WIDE field imaging surveys coming online in the next years

- Cosmology end to end tests of the standard model, dark energy/modified GR, neutrino mass/hierarchy
- Galaxy evolution environmental studies relating to cosmic web, baryon cycle studies
- Galactic archeology kinematics for DM probes, chemical tagging for assembly histories
- Transient universe classifications of SN types, exploring phase-space for new classes of variables
- Solar system trans-Neptunian studies





Ground-Based

4m class funded

8-10m class funded

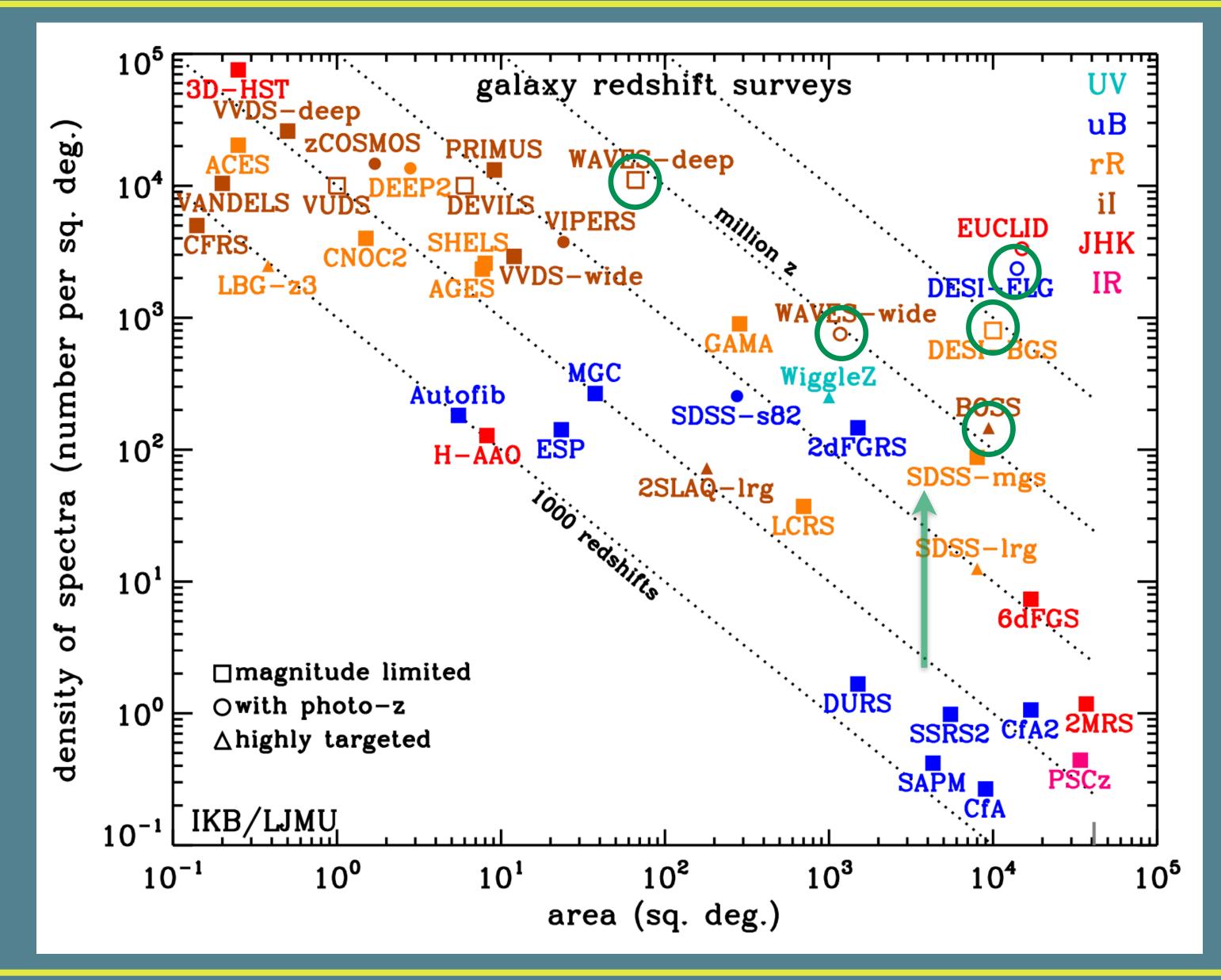
Proposed & unfunded

Instrument/Telescope	Collecting Area m ²	Field of view deg ²	Multiplex	R
4MOST	10.7	4.00	1400	6500-2
Mayall 4m / DESI 🗹	11.4	7.08	5000	2000-5
WHT / Weave	13.0	3.14	1000	5000-2
Subaru / PFS	52.8	1.25	2400	3000-5
VLT / MOONS	52.8	0.14	500	4000-6 9000-2
Keck / DEIMOS ?	76.0	0.015	150	
Megamapper @ GMT	28.0	7.06	20,000	
Keck / FOBOS	76.0	0.087	1800	
MSE @ CFHT	78.5	1.52	4000	
ESO Spectel	113.1	4.90	5000	

THE ERA OF LARGE SURVEYS



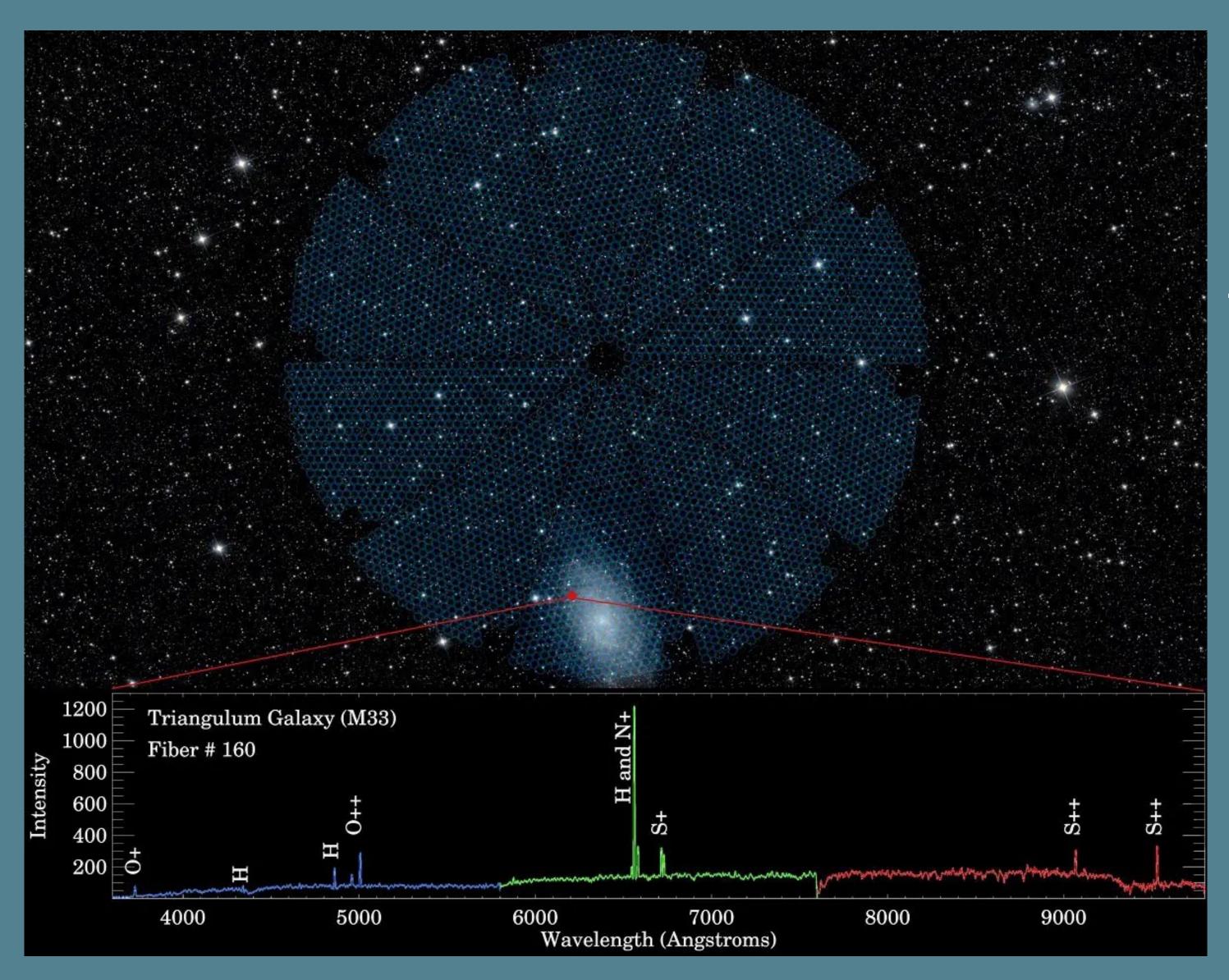




THE ERA OF LARGE SURVEYS

I. K. Baldry

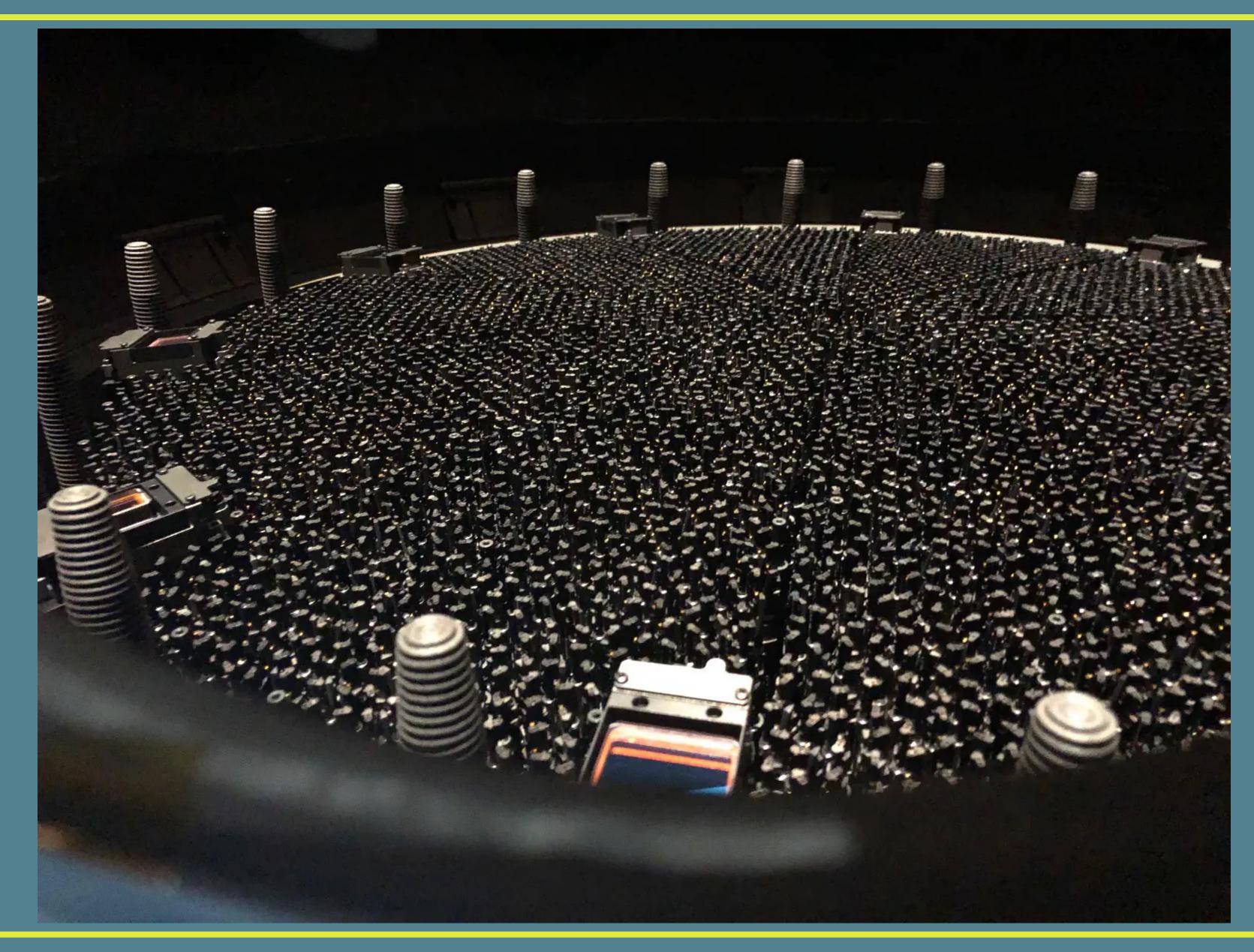
DARK ENERGY SPECTROSCOPIC INSTRUMENT (DESI)



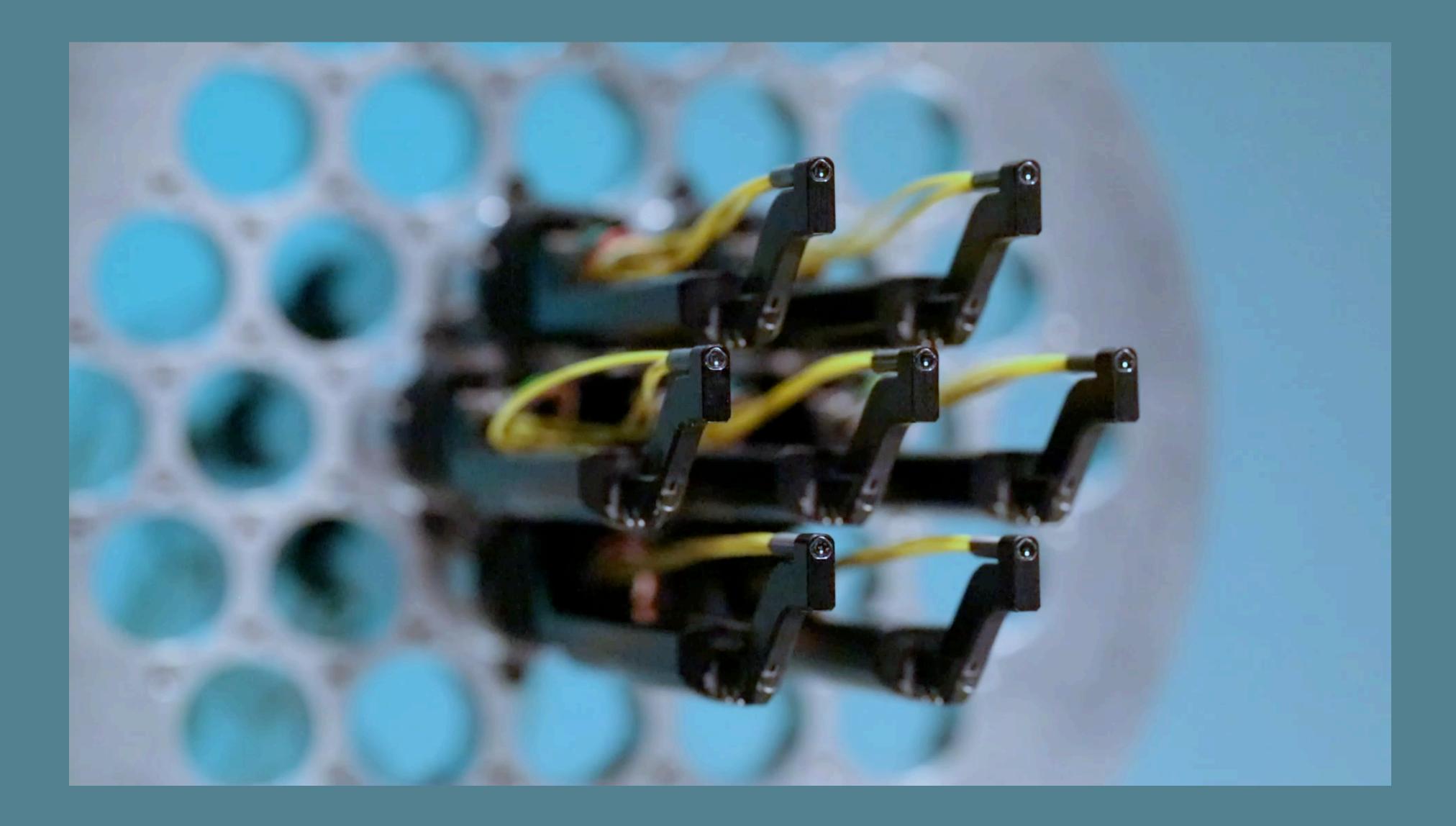
DESI Early Data Release contains 2M spectra from the Survey Validation phase of galaxies, quasars and stars

Targets have been selected from the Legacy Surveys (DECaLS, BASS and MzLS)

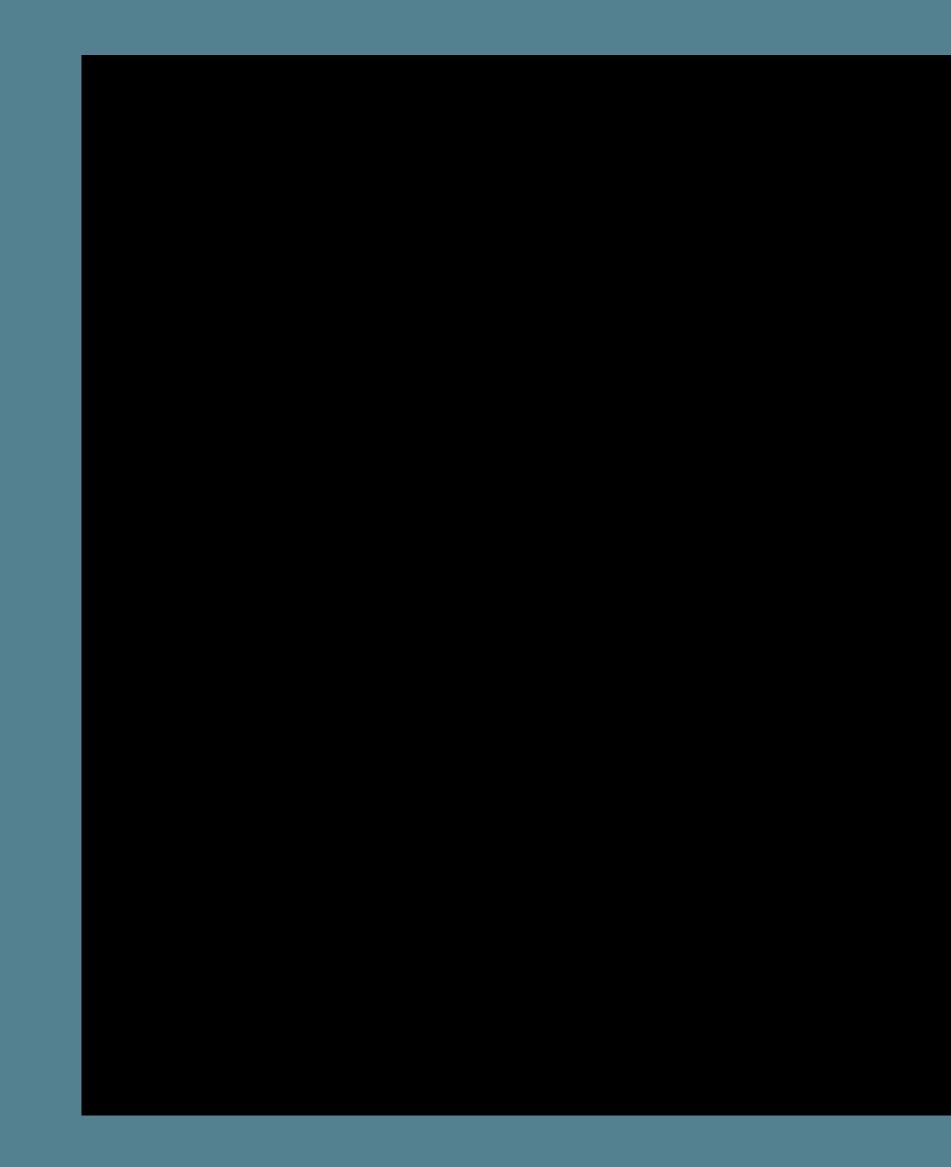
DARK ENERGY SPECTROSCOPIC INSTRUMENT (DESI)



SDSS-V FIBERS FOR APOGEE-2

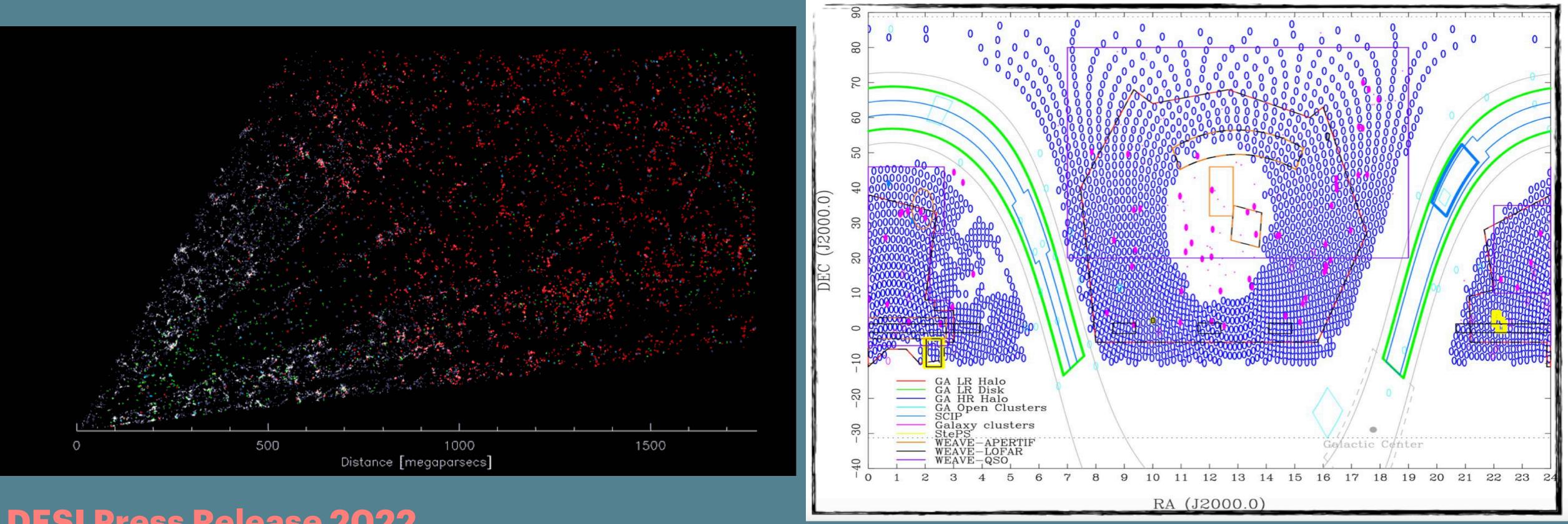


WHT ENHANCED AREA VELOCITY EXPLORER (WEAVE)



THE ERA OF LARGE SURVEYS: HIGHLIGHTS

4m-class: DESI (Mayall) / WEAVES (WHT)

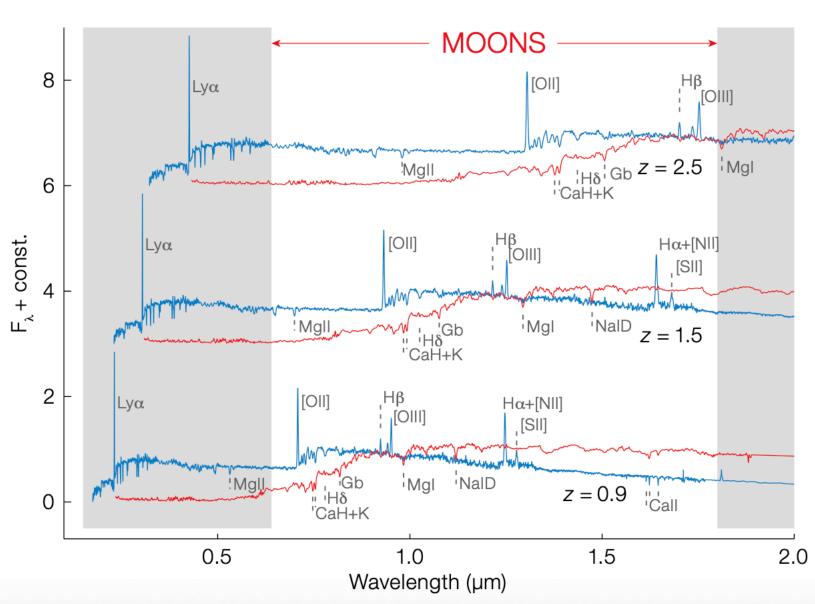


DESI Press Release 2022

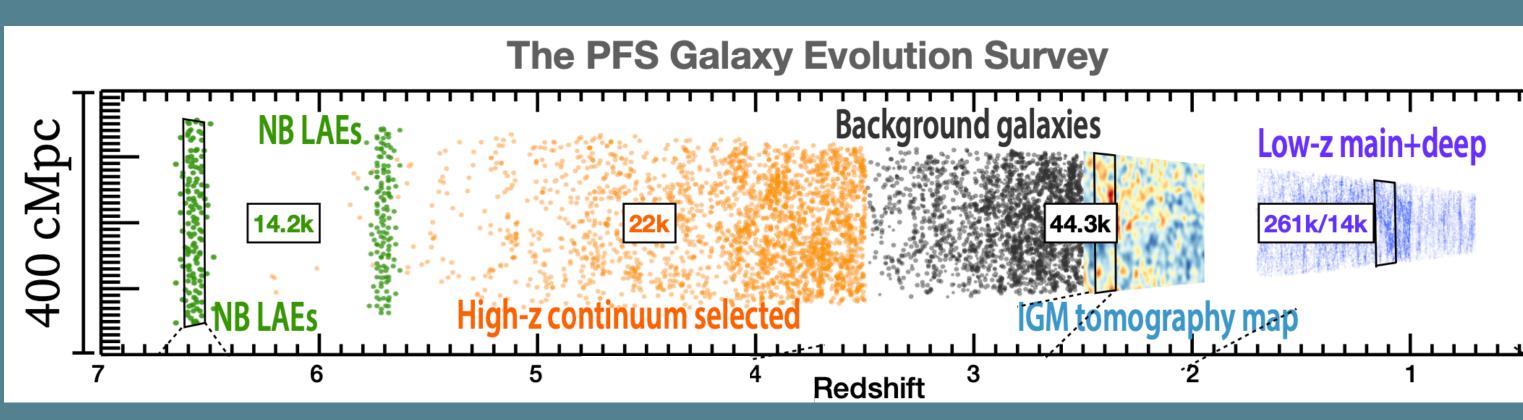
By S. Trager

THE ERA OF LARGE SURVEYS: HIGHLIGHTS

8m-class: MOONS (VLT) / PFS (Subaru)

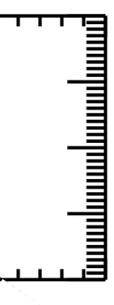


Wavelength range (µm)



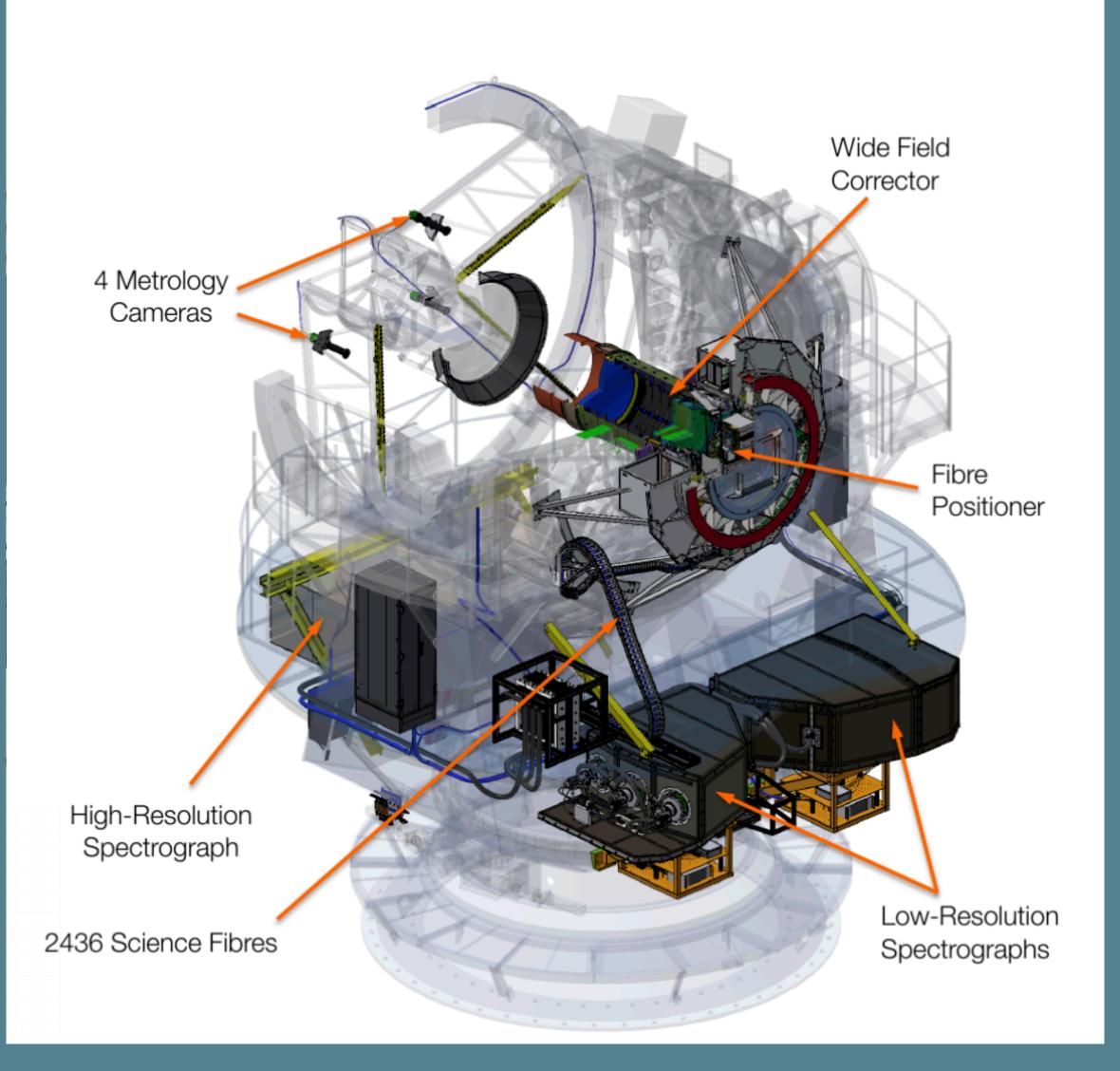
Maiolino+2020

Greene+2022





THE 4M MULTI-OBJECT-SPECTROSCOPIC-TELESCOPES (4MOST)



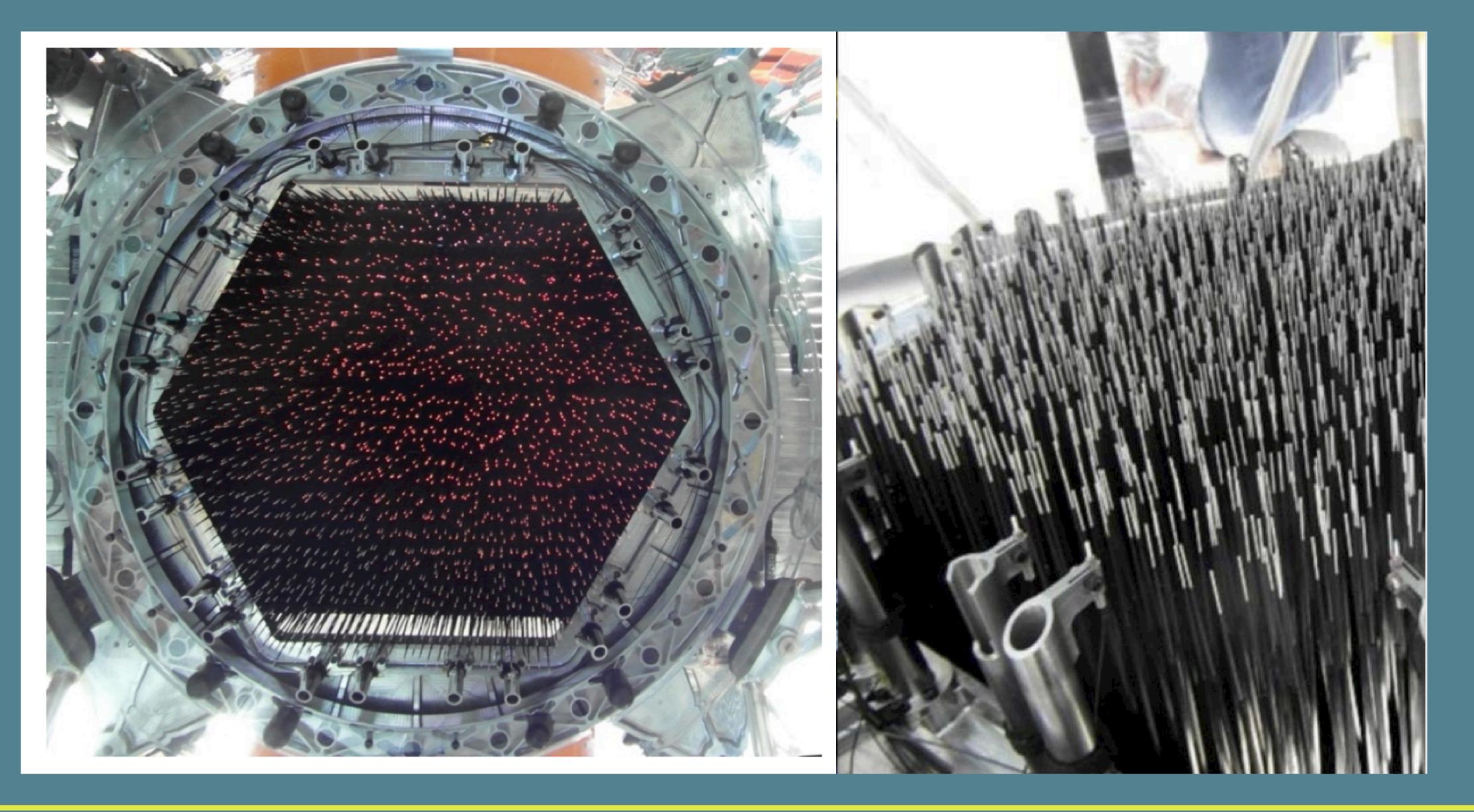
Low Resolution Spectrographs (2x)	Passband: 370-950nm $R > \lambda \times 10$ for 400nm $< 500nm$ R > 6000 for 500nm $< 885nmVelocity Accuracy <1 km/s1624 fibres in total$
High Resolution Spectrograph	Passbands: 392.6-435.5, 516-573, 610-679 nm <i>R</i> > 18,500 Velocity Accuracy <1km/s 812 fibres

In a 5-year survey 4MOST of survey observations

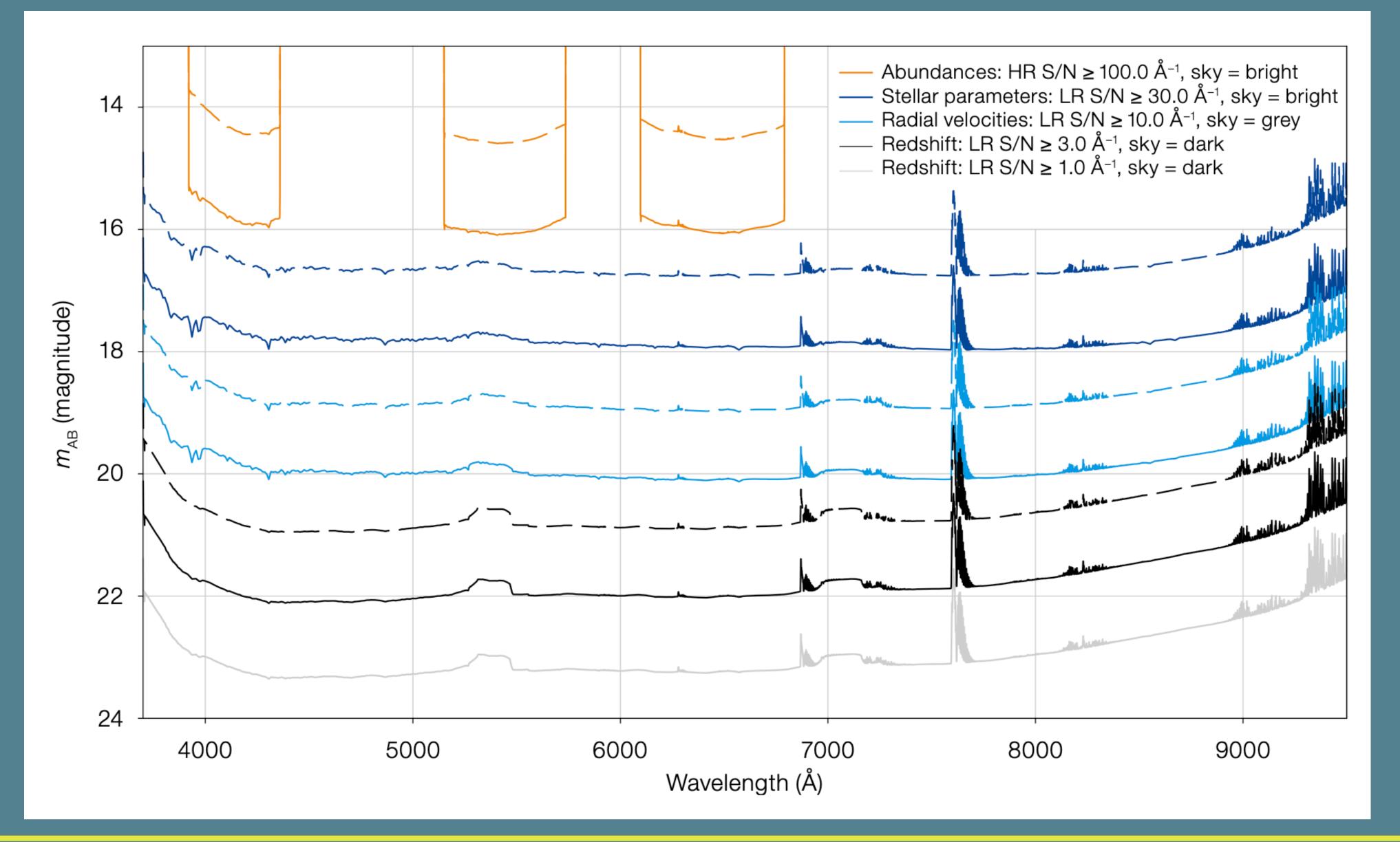
- Cover at least twice an area of =16,000 degree² (goal >20,000)
- Obtain >15 million (goal >25 million) spectra at resolution R~5000
- Obtain >1 million (goal >2 million) spectra at resolution R~20,000.

VISTA Telescope @ Paranal





4MOST



4MOST

Consortium Surveys (70%)

- Milky Way Halo LR Survey
- Milky Way Halo HR Survey
- Milky Way Disk and Bulge LR Survey
- Milky Way Disk and Bulge HR Survey
- Galaxy Clusters Survey
- AGN Survey
- Galaxy Evolution Survey (WAVES)
- Cosmology Redshift Survey
- Magellanic Clouds Survey
- Time-Domain Extragalactic Survey (TIDES) Sullivan (Southampton)

Irwin (IoA), Helmi (RuG) Christlieb (ZAH) Chiappini, Minchev, Starkenburg (AIP) Bensby (LU), Bergemann (MPIA) Finoguenov (MPE) Merloni (MPE) Driver (UWA), Liske (UHH) Kitaura (AIP), Richard (CRAL), Kneib (EPFL) Cioni (AIP)





Community Surveys (30%)

Toloza, O. et al. – The White Dwarf Binary Survey (WDB) Sacco, G. G. et al. – The 4MOST Survey of Young Stars (4SYS) Ibata, R. et al. – 4MOST Gaia RR Lyrae Survey (4GRoundS) Lucatello, S. et al. – Stellar Clusters in 4MOST **Pawlak, M. et al.** – Spectroscopic Discovery of Binaries with Dormant Black Holes Skúladóttir, Á. et al. – The 4MOST Survey of Dwarf Galaxies and their Stellar Streams (4DWARFS)

Iovino, A. et al. – Stellar Population Survey Using 4MOST (4MOST-StePS) Duncan, K. et al. – Optical, Radio Continuum and HI Deep Spectroscopic Survey (ORCHIDSS)

Gruen, D. et al. – 4MOST Complete Calibration of the Colour-Redshift Relation (4C3R2)

Haines, C. et al. - CHANCES: A CHileAN Cluster galaxy Evolution Survey Bauer, F. E. et al. – Chilean AGN/Galaxy Extragalactic Survey (ChANGES) **Krogager, J.-K. et al.** – The 4MOST–Gaia Purely Astrometric Quasar Survey (4G-PAQS)

Peroux, C. et al. – Transform our Understanding of the Baryon Cycle with High-Resolution Quasar Spectroscopy (ByCycle) Taylor, E. N. et al. – The 4MOST Hemisphere Survey of the Nearby Universe (4HS) Collett, E. T. et al. – The 4MOST Strong Lensing Spectroscopic Legacy Survey (4SLSLS)

4MOST

Messenger Nº190



~ 10% of the 30% of 4MOST Comm Survey time: ~ 2 Mhrs of low-resolution spectroscopy for a representative sample of (mostly) variability and SED selected AGN

Variability selection: Currently ZTF + CLQ In the future: Rubin LSST survey We are using advanced AI techniques (see the ALeRCE Explorer: https://alerce.online)

SED selection: **Optical-NIR-MIR SED modeling to look for warm dust**

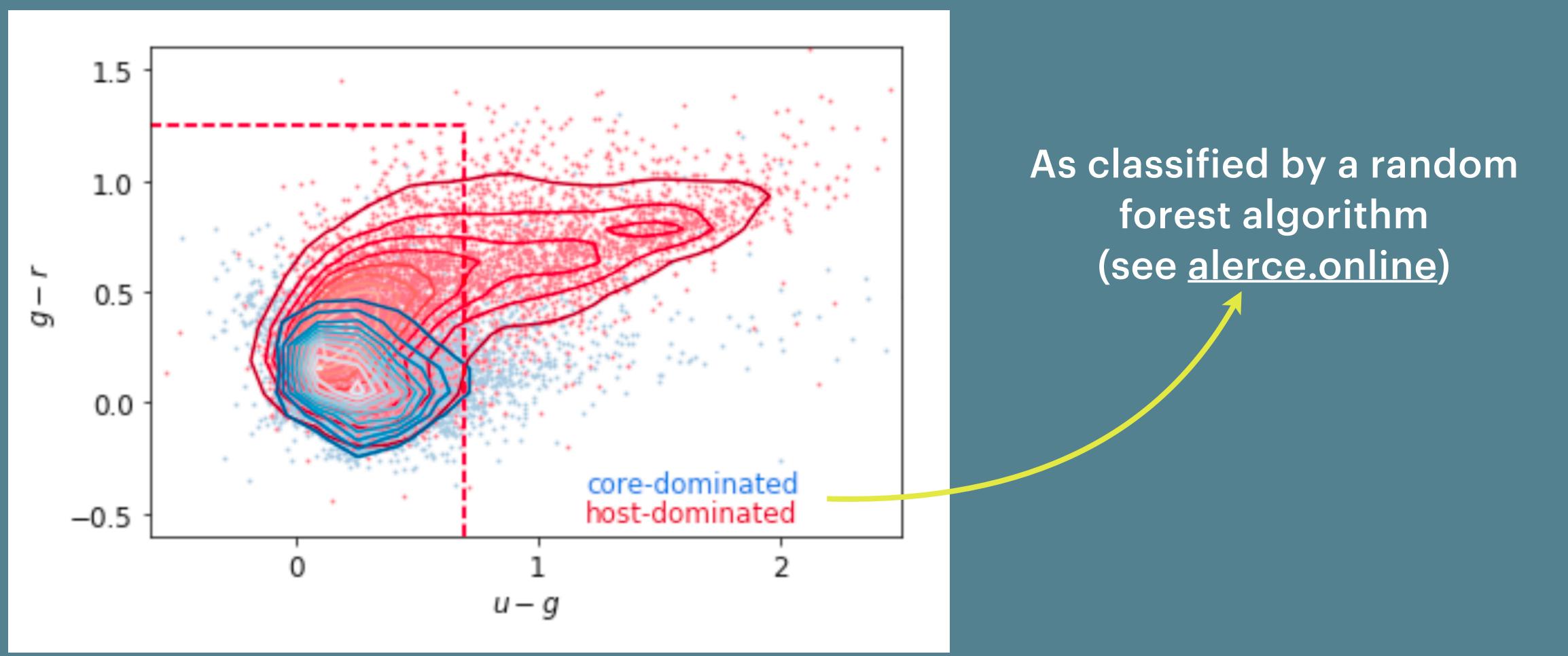
+ Changing Type / State AGN, TEDs, Lenses, Intervening QALs

CHILEAN AGN / GALAXIES EXTRAGALACTIC SURVEY (CHANGES)





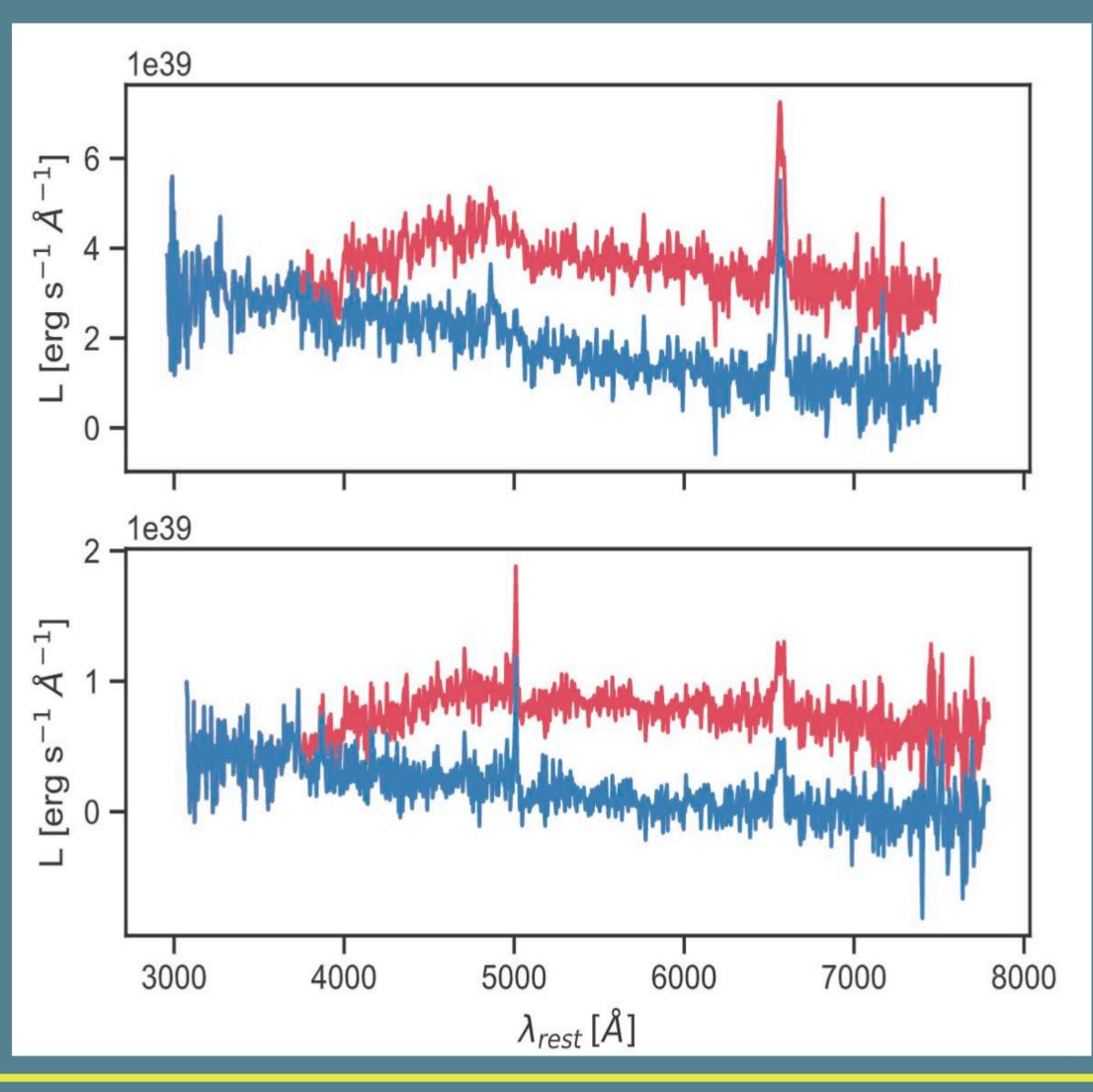
Variability selection:







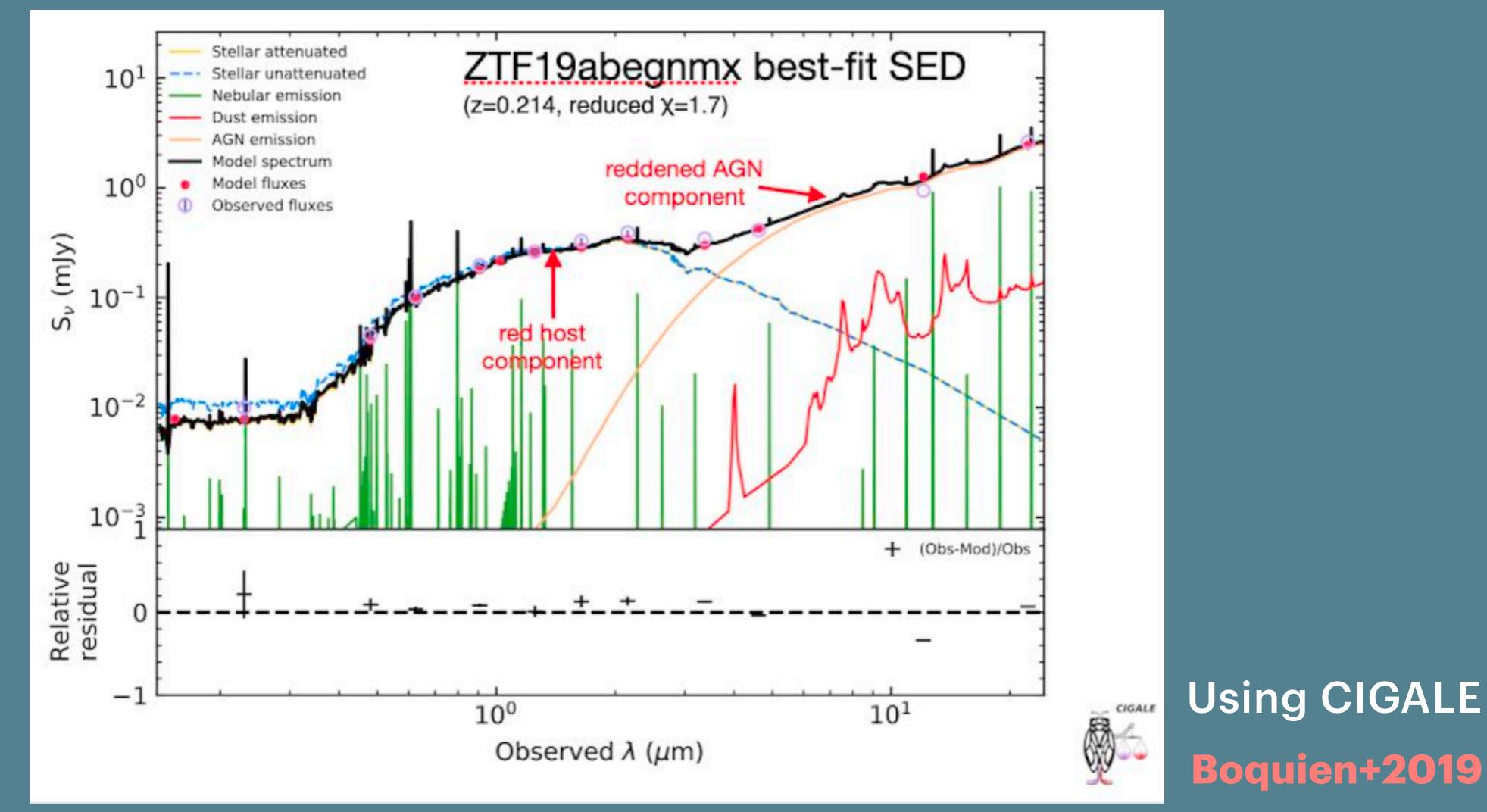
Variability selection:



Sánchez-Saez+2019

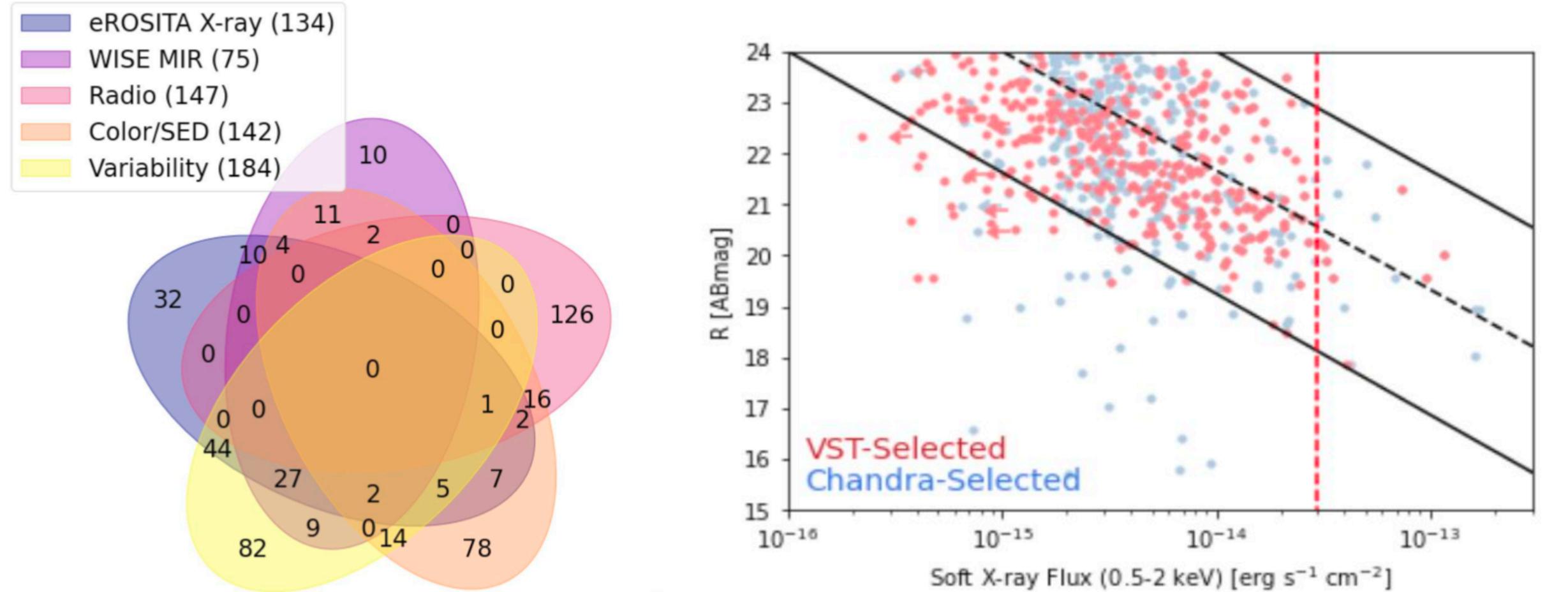


SED selection:





Overlap with eROSITA GTO AGN Survey



WIDE AGN Distribution (R<22.5) cut: 482 AGN/deg²

Main Measurables:

- Redshifts
- **Emission line measurements**
- **Detection of broad components**
- Host Galaxy characterization
- Variations in SED / line emission profiles

Main Deliverables:

BH accretion rate densities, evolution and host synergies for moderately accreting AGN ($10^{-4} < L/LEdd < 10^{-1}$) that comprises ~ 50–80% of the estimated total mass accretion onto BHs in type 1 and mildly obscured AGN that strongly complements other 4MOST AGN samples.